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## Probabilistic Seismic Source Inversion of the 1886 Charleston, South Carolina, Earthquake from Macroseismic Evidence: A Major Updating

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## 5 ABSTRACT

6 The source of the 1886 Charleston, South Carolina earthquake influences the computed seismic hazard of 7 the Southeastern U.S. and thus impacts public policy and engineering practice. However, because the 1886 8 rupture predated seismic instruments, its source is highly uncertain. This study presents probabilistic 9 seismic-source inversions of the Charleston earthquake from liquefaction evidence and historical intensity 10 reports. Using the latest predictive models and a novel inversion approach, we seek to constrain the 11 magnitude, location, and orientation of the 1886 rupture. Probability distributions of rupture magnitude are 12 conditioned on both the "Woodstock Fault" - a commonly inferred source of the 1886 event - and on an 13 unknown source, wherein the uncertainties of fault location and orientation are considered. These 14 distributions are compared to the M<sub>w</sub>6.7-M<sub>w</sub>7.5 distribution adopted by the U.S. National Seismic Hazard Model Project (NSHMP). Collectively, the results do not provide strong support for the hypothesized 15 16 Woodstock Fault. This is not to say the Woodstock Fault does not exist, but rather, that the position of the 17 1886 source model cannot be constrained by the data and models studied herein, given the large 18 uncertainties inherent to each. While this is at odds with the underlying assumption of many prior studies, 19 the results nonetheless generally uphold the magnitude distribution assumed by the NSHMP. The largest 20 uncertainties inherent to this distribution are identified and could be diminished in the future. Finally, we 21 note that the inversion methodology used here is not specific to any region, or to certain types of evidence, 22 but can be applied to any seismic zone and to any co-seismic response. This methodology allows for 23 uncertainty to be accounted for in a more complete and transparent manner when inverting seismic source 24 parameters from macroseismic data. Of course, any limitations, biases, or unmodeled uncertainties inherent 25 to these data must be understood, and their implications acknowledged, as further discussed herein.

26 Keywords: seismic hazard, macroseismic intensity, soil liquefaction, inverse analysis

# 27 1. Introduction

Computed seismic hazards are especially uncertain in regions of infrequent seismicity, where the return periods of moderate-to-large earthquakes may exceed the historic observational period. To reduce this uncertainty, engineering geologists routinely perform forensic analyses of the macroseismic evidence (e.g., liquefaction, landslides, intensity reports) produced by prehistoric and pre-instrumental earthquakes (e.g.,

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among many, Obermeier et al., 1998; Schneider et al., 2001; Green et al., 2005; Kuhn, 2005; Obermeier et
al., 2005; Olson et al., 2005; Rodriguez-Marek and Ciani, 2008; Maurer et al., 2015a; Gheibi and Gassman,
2016; Yousuf et al., 2021; Chung et al., 2021; Rasanen et al., 2021; Bwambale et al., 2022). The goal of
these analyses, in effect, is to constrain the seismic-source parameters of paleoearthquakes, such that these
parameters may be input to probabilistic seismic hazard analyses. It follows that computed seismic hazards
are, in some regions, heavily influenced by analyses of macroseismic evidence. The South Carolina Coastal
Plain is one such region and is strongly influenced by interpretations of the 1886 Charleston earthquake.

39 The 1886 event induced widespread soil liquefaction across the Coastal Plain (Amick et al., 1990), 40 damaged structures in multiple U.S. states, including most structures in Charleston (Dutton, 1889; Wong et 41 al., 2005), produced perceptible shaking over 1500 km away in Canada (Bakun et al., 2002), and was larger 42 in magnitude than any earthquake to since occur in the Southeastern U.S. The source of the Charleston 43 event is thus a major seismic hazard for the region. A 2005 study, for example, predicted that a repeat of the 1886 event would cause 900 deaths, 44,000 injuries, and economic losses of \$20 billion in South 44 45 Carolina alone (Wong et al., 2005). In turn, the Charleston source controls the computed seismic hazard for much of the Southeastern U.S., particularly for long-period structures (Petersen et al., 2020), and thus 46 47 impacts building codes, governing policies, and engineering practice. However, because the 1886 rupture 48 predated seismic instruments and did not manifest at the surface, its exact location and magnitude remain 49 uncertain, as do the regional amplitudes of resultant ground motions. To constrain these unknowns, and 50 thus prepare for a similar event, numerous researchers have studied macroseismic evidence, as summarized 51 in Table 1. Published confidence intervals (CIs) of the 1886 magnitude range from  $M_w 6.4$  to  $M_w 7.8$ , as 52 interpreted from intensity reports (e.g., Bakun and Hopper, 2004; Cramer and Boyd, 2014) and liquefaction 53 evidence (e.g., Martin and Clough, 1994; Hayati and Andrus, 2008). Liquefaction features also suggest a 54 history of recurrent earthquakes in the region extending back 6,000 years (Talwani and Schaeffer, 2001) 55 with wide-ranging magnitude estimates of M<sub>w</sub>5.1 to M<sub>w</sub>7.8 (e.g., Hu et al., 2002; Gheibi et al., 2020). 56 Considering the existing literature, Petersen et al. (2014, 2020) assigned to the Charleston seismic zone a magnitude probability distribution that ranged from M<sub>w</sub>6.7 to M<sub>w</sub>7.5 in the most recent U.S. National 57 58 Seismic Hazard Model Project (NSHMP) maps.

59 While much has been learned about the 1886 Charleston earthquake, prior analyses of the macroseismic 60 data (i.e., intensity reports, soil liquefaction) have several limitations. First, the analysis of this data has 61 multiple uncertainties, yet existing studies tend either to be deterministic or to account for uncertainties 62 informally. That is, they generally provide either a median estimate of the rupture magnitude or uncertainty 63 bounds that are nominal in nature. It is often unclear what the bounds are, exactly, and which uncertainties 64 are, and are not, accounted for. Second, most analyses assume that the 1886 event was caused by a particular 65 fault (i.e., the "Woodstock Fault") with known characteristics, even though the fault(s) responsible for the 66 event are debated and the characteristics of the Woodstock Fault are uncertain. The feasibility of the data 67 to constrain the source model has arguably not been fully explored, given that nearly all studies provide a 68 magnitude estimate conditioned on a single fault and do not investigate the uncertainty of this assumption. Third, the inverse analysis of intensity and liquefaction data requires a series of models for predicting these 69 70 phenomena. Regionally distributed ground motions must be predicted, conditioned on a hypothetical 71 source, to include site-response effects at the locations of study. The probability of field observations (i.e., 72 the observed intensity or liquefaction response) must then be computed, conditioned on the expected ground 73 motions. In this regard, major modeling advances have recently been made. The NGA-East project (Goulet 74 et al., 2018) resulted in the most advanced understanding of Eastern North America (ENA) ground motions 75 and site response (Harmon et al., 2019) to date. Models for correlating ground motions to macroseismic 76 intensities, including ENA-specific relationships, have been updated (e.g., Cramer, 2020). And models for 77 predicting the probability of liquefaction surface expression have been trained using all liquefaction case 78 histories globally compiled to date (Geyin and Maurer, 2020).

**Table 1.** Prior estimates of the 1886 Charleston earthquake magnitude; estimates are in moment magnitude  $(M_w)$  and ranges are at the 95% confidence level, unless noted otherwise  $(m_b = body wave magnitude; M_s = surface wave magnitude).$ 

Study	Study Type	Magnitude	
Bollinger (1977)	MMI	6.8-7.1 (m <sub>b</sub> )*	
Nuttli et al. (1986)	MMI	6.7 (m <sub>b</sub> ), 7.7 (M <sub>s</sub> ) <sup>†</sup>	
Martin and Clough (1994)	Liquefaction	7.0-7.5 <sup>†</sup>	
Johnston (1996)	MMI	6.8-7.8	
Bakun and Hopper (2004)	MMI	6.4-7.2	
Heidari and Andrus (2010)	Liquefaction	$6.8-7.0^{+}$	
Cramer and Boyd (2014)	MMI	6.7-7.3	

\* Upper bound magnitude estimate (range is not at the 95% confidence level)

<sup>†</sup> Magnitude range is either not given or is not at the 95% confidence level

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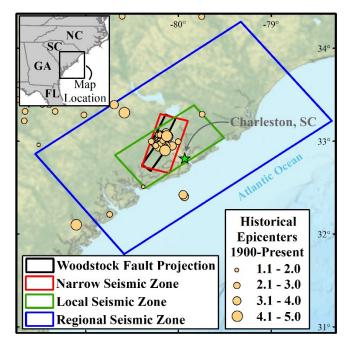
80 Accordingly, this study presents probabilistic seismic-source inversions of the 1886 Charleston earthquake from historical intensity reports and liquefaction evidence. Each is studied using a novel 81 82 approach wherein the above shortcomings are addressed directly. With this approach, the likelihood of a 83 rupture with some location, geometry, and magnitude to produce a set of field observations (observed 84 intensities or liquefaction responses) is computed. Repeating for enumerable hypothetical faults results in 85 a regional scale understanding of the likely source parameters, to the degree the observational data permits. Probability distributions of earthquake magnitude, conditioned on both an unknown source and on the 86 Woodstock Fault, are computed and compared to that used to develop the NSHMP maps (Petersen et al., 87 88 2014, 2020). In the following, prior studies of the 1886 macroseismic data are summarized. An overview

of the analysis methodology is then presented, followed by implementation details. Lastly, the
macroseismic data are analyzed and a variety of results are presented and discussed.

### 91 2. Prior analyses of 1886 macroseismic evidence

# 92 2.1 Analyses of MMI data

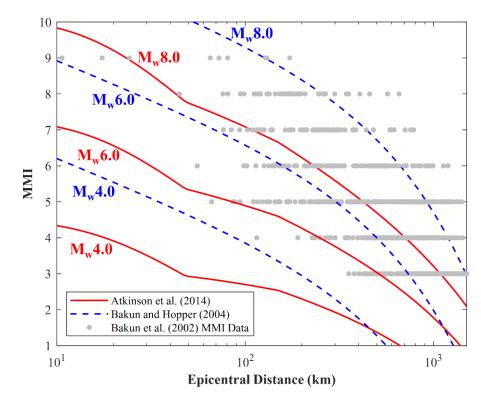
Following the 1886 earthquake, Dutton (1889) compiled intensity reports throughout ENA and developed 93 94 isoseismal maps based on the Rossi-Forel intensity scale. Researchers have since reinterpreted these reports 95 to the Modified Mercalli Intensity (MMI) scale and analyzed them to infer seismic parameters. Bakun et 96 al. (2002), for example, compiled 1,034 MMI observations from these and other original reports. Prior 97 analyses of the 1886 MMI data have typically used intensity-prediction equations (IPEs), which predict intensity as a function of rupture magnitude and site-to-source distance. Assuming some source location 98 99 and adopting an IPE, researchers have constrained the causative M<sub>w</sub> which best fits the MMI data (e.g., Bollinger, 1977; Nuttli et al., 1986; Johnston, 1996; Bakun and Hopper, 2004) as summarized in Table 1. 100 101 Bakun and Hopper (2004), for example, developed an ENA-specific IPE, applied it to the Bakun et al. (2002) MMI data, and reported a magnitude of M<sub>w</sub>6.9 (M<sub>w</sub>6.4-7.2 at the 95% confidence level). Using 102 103 different methods than prior researchers, Cramer and Boyd (2014) compared the mean MMI of the Bakun 104 et al. (2002) dataset against those from two reference events in similar tectonic settings (M<sub>w</sub>7.2 1929 Grand 105 Banks, Canada and M<sub>w</sub>7.6 2001 Bhuj, India) over a site-to-source distance of 600-1200 km. With this 106 approach, Cramer and Boyd (2014) estimated a median magnitude of  $M_w$ 7.0 with uncertainty of  $\pm 0.3 M_w$ . Collectively, existing studies of the MMI data have reported estimates of  $M_w 6.4$  to  $M_w 7.8$ . In producing 107 108 such estimates, these studies have generally assumed that the source was epicentrally located in the vicinity of what is typically called the Woodstock Fault, an inferred N striking, W dipping fault ~25 km NW of 109 110 Charleston (e.g., Durá-Gómez and Talwani, 2009a,b; Chapman et al., 2016). The fault's surface projection, 111 as hypothesized by Durá-Gómez and Talwani (2009a,b), is mapped in Figure 1. While most studies based on geophysical investigations or modern seismological data have supported this proposed alignment - at 112 least in a general sense (e.g., Pratt et al., 2022) – dramatically different hypotheses for the 1886 earthquake 113 have also been proposed. Marple and Hurd (2020), for example, recently suggested that the 40-km long 114 115 "Deer Park lineament," which is oriented roughly E-W, may have been responsible.



**Figure 1.** Woodstock Fault projection (Durá-Gómez and Talwani, 2009a,b) and the zonal weighting scheme assigned to the Charleston Seismic Zone in the most recent U.S. National Seismic Hazard Model Project (NSHMP) maps (Petersen et al., 2014, 2020).

Also shown in Figure 1 is the zonal weighting scheme assigned to the Charleston Seismic Zone by the 116 NSHMP (Petersen et al., 2014, 2020). In this scheme, a "Narrow" zone with weight of 0.3 delineates the 117 hypothesized Woodstock Fault while accounting for uncertainties in its position and branching structure. 118 119 The "Local" and "Regional" zones, with respective weights of 0.5 and 0.2, collectively extend offshore and across the extents of the South Carolina Coastal Plain, thereby encompassing more distal faults and 120 liquefaction features that have not been tied to the 1886 event. Each zone is assigned the same  $M_w 6.7$  to 121 M<sub>w</sub>7.5 probability distribution by Petersen et al. (2014, 2020). This weighting scheme was adopted from 122 the Central and Eastern United States Seismic Source Characterization for Nuclear Facilities Project 123 124 (Coppersmith et al., 2012), who concluded: "Neither the 1886 nor the prehistoric (i.e., pre-1886) 125 earthquakes in the Charleston area can be definitively attributed to any specific fault or fault zone at the 126 present time." In arriving at this conclusion, Coppersmith et al. (2012) noted: "the Charleston region is 127 associated with a pattern of observed seismicity that is not particularly remarkable for drawing attention to the location of the 1886 earthquake." While it should be emphasized that the NSHMP weighting scheme 128 does not describe the uncertainty of the 1886 earthquake specifically, it does reflect the overall uncertainty 129 of moderate-to-large "1886-like" earthquakes in the region. Thus, while prior studies have generally 130 assumed that the 1886 source was in the "Narrow" zone, a scientific consensus has not been reached. 131

132 As a precursor to other analyses presented herein, the Bakun and Hopper (2004) IPE was first 133 reimplemented on the Bakun et al. (2002) dataset and the same  $M_w 6.9$  estimate as Bakun and Hopper (2004) 134 was obtained, indicating that the approach and dataset were correctly reproduced. Next, this approach was updated using the newest ENA IPE (Atkinson et al., 2014). Assuming the same source location and studying 135 MMI data within 1000 km (the applicable distance of the Atkinson et al. (2014) IPE) and within all distances 136 (to mirror Bakun and Hopper, 2004), median estimates of M<sub>w</sub>8.0 and M<sub>w</sub>8.2 were respectively obtained. 137 138 The cause of the discrepant results obtained using Bakun and Hopper (2004) vs. Atkinson et al. (2014) can be seen in Figure 2, where both IPEs are plotted for three values of  $M_w$ . For a given  $M_w$  and epicentral 139 140 distance, the Atkinson et al. (2014) model tends to predict a lesser MMI, indicating that a larger earthquake 141 magnitude (i.e.,  $M_w 8.0 - M_w 8.2$ ) is needed to produce the same set of MMI observations.



**Figure 2.** ENA-specific intensity prediction equations (IPEs) proposed by Bakun and Hopper (2004) and Atkinson et al. (2014) considering three values of  $M_w$ .

While it may appear, per the latest ENA IPE, that the 1886 rupture was much larger than previously thought, there are limitations that give rise to the work that follows. Most notably, perhaps, is that existing ENA IPEs do not allow for consideration of site effects when predicting MMI. The Atkinson et al. (2014) IPE, for example, is intended for site class C conditions (i.e., stiff soils). Accordingly, if some of the 1886 MMI observations were on softer sites (e.g., in river valleys or along waterways, which seems likely), then the estimate of M<sub>w</sub>8.0-M<sub>w</sub>8.2 obtained via the Atkinson et al. (2014) IPE could require significant reduction. 148 Moreover, the MMI studies in Table 1: (i) predate the latest knowledge of ENA ground motions and site 149 response (Goulet et al., 2018; Harmon et al., 2019); and (ii) do not rigorously account for uncertainty. In 150 this study, probabilistic site-adjusted ground motion intensity measures (IMs) will be explicitly predicted using 17 ENA ground-motion models (GMMs) (whereas ground motions were only implicitly predicted in 151 152 prior studies). In turn, MMI values will be probabilistically predicted at study sites using the latest IM-MMI 153 models (e.g., Cramer, 2020). These predictions will be repeated for a multitude of fault locations, 154 orientations, and magnitudes to compute the likelihood that each source would produce the 1886 MMI 155 observations of Bakun et al. (2002). This approach, which will be subsequently presented in detail, 156 incorporates ground-motion IM uncertainty, IM-MMI uncertainty, and source location uncertainty to 157 produce a probability distribution of rupture magnitude. It must be noted that MMI data is also subject to 158 measurement uncertainty and reporting bias (e.g., Hough et al., 2000; Cramer and Boyd, 2014). While site-159 specific measurement uncertainties, correction factors, or weighting schemes could be accommodated, a thorough reinterpretation of the more than 1000 original intensity reports would be required, to include 160 161 possible reassignment of MMI values in the Bakun et al. (2002) dataset and development of observationspecific uncertainties. In the current effort, however, we study the existing data directly, treat measurement 162 163 uncertainty in a simple manner, and assign all observations equal weight. It should be noted that all prior 164 studies of the 1886 intensities have also used these data. Thus, while our methodology has important 165 advantages over prior efforts (e.g., the capacity to probabilistically constrain the rupture location), a future 166 study might further benefit from rigorous reinterpretation of the original intensity reports.

## 167 *2.2 Analyses using liquefaction*

168 Paleoliquefaction evidence suggests that at least seven moderate-to-large earthquakes have impacted the 169 South Carolina Coastal Plain in the last 6,000 years (e.g., Gohn et al., 1984; Obermeier et al., 1985; 170 Obermeier et al., 1987; Talwani and Schaeffer, 2001). Three of these events are interpreted to have a source in the vicinity of Charleston and a recurrence rate of ~500 years (Talwani and Schaeffer, 2001). The 1886 171 172 event was the most recent of these to generate liquefaction and the only instance in which liquefaction was well documented as it occurred. Surface manifestations (e.g., ejecta, ground cracks) were mapped by Earle 173 174 Sloan, among others, and compiled by Dutton (1889). Additional liquefaction evidence was subsequently discovered during trenching investigations (e.g., Obermeier et al., 1985; Talwani and Cox, 1985). 175 Collectively, this evidence has been analyzed to determine the magnitude of the earthquake that caused it. 176 177 Liquefaction models (e.g., Green et al., 2019; Maurer et al., 2015b) conventionally predict the future 178 triggering and surface manifestation of liquefaction, given in-situ geotechnical test data and some seismic 179 loading. In an inverse analysis, these models are used in reverse to constrain the seismic loading that would, 180 and would not, produce the observed response. By comparing this loading to that forward predicted by a

181 GMM for an assumed source, the magnitude of that source may be constrained. Studying primarily standard

182 penetration test (SPT) data from sites of interest, Martin and Clough (1994) carried out such an analysis

183 with the Seed et al. (1984) SPT-based liquefaction triggering model and the Ishihara (1985) liquefaction

184 manifestation model. Assuming a seismic source at the centroid of reported intensity (roughly consistent

185 with the hypothesized Woodstock Fault), adopting GMMs then available (e.g., Chapman et al., 1989), and

- 186 employing considerable judgement, Martin and Clough (1994) estimated that an M<sub>w</sub>7.0-M<sub>w</sub>7.5 event could
- 187 produce liquefaction consistent with that observed.
- 188 Studying cone penetration test (CPT) data, Hayati and Andrus (2008) used the Robertson and Wride 189 (1998) CPT-based triggering model and the Iwasaki et al. (1978) manifestation model to estimate a magnitude of  $M_w 6.8$ - $M_w 7.3$ . Because triggering models are trained almost exclusively with case-history 190 data from Holocene deposits, "aging correction" factors (or "deposit resistance" corrections) have been 191 192 proposed when applying triggering models in older soils. Specifically, it has been argued that aging effects, or increases in the cyclic strength of soils over time, may be resolved into gains measurable by large strain 193 194 penetration tests and gains influenced by soil fabric phenomena undetectable at larger strain (e.g., Maurer et al., 2014). Thus, penetration resistance may correlate to liquefaction resistance differently in Pleistocene 195 196 soils than in Holocene soils. Accordingly, Hayati and Andrus (2008) employed aging correction factors in 197 select geologic units. In producing their estimate of M<sub>w</sub>6.8-M<sub>w</sub>7.3, Hayati and Andrus (2008) assumed that 198 the Woodstock fault was the source and that a peak ground acceleration (PGA) of 0.3 g occurred throughout 199 Charleston, citing previous ground-motion predictions (e.g., Silva et al., 2003). Using this methodology, 200 they constrained the magnitude range for which predictions from liquefaction models matched observations of response. Heidari and Andrus (2010) used a similar methodology as Hayati and Andrus (2008) but 201 202 applied the updated aging correction factors of Hayati and Andrus (2009) to obtain an estimate of  $M_w 6.8$ - $M_w$ 7.0. In addition to these studies, researchers have studied paleolique faction interpreted to be from older, 203 possibly similar events. Gheibi et al. (2020), for example, studied evidence induced by a "Charleston 204 Source" approximately ~550 and ~5,000 YBP. Assuming the source to be the Woodstock fault, they 205 computed respective minimum magnitudes of M<sub>w</sub>6.6-M<sub>w</sub>7.2 and M<sub>w</sub>6.2-M<sub>w</sub>6.7 for these two events, where 206 207 the uncertainty stems from which GMM is adopted to predict median ground motions.
- Like prior studies of the MMI data, those of liquefaction evidence have greatly improved knowledge of the regional seismic hazard, but also have limitations that motivate the present study. In brief, existing studies: (i) predate both the NGA East Project (Goulet et al., 2018; Harmon et al., 2019) and the latest liquefaction models trained on all globally available data (Geyin and Maurer, 2020); and (ii) do not rigorously account for uncertainty. In this regard, prior studies do not account for source-model uncertainty (i.e., they assume a single seismic source), do not account for the uncertainty of ground motions conditioned on that source, and do consider the prediction of liquefaction in any probabilistic sense. In general,

published uncertainties of the 1886 rupture magnitude, whether derived from MMI or liquefaction data, are arguably nominal in nature. In some studies, for example, only a single uncertainty is considered, such as the epistemic uncertainty of which deterministic model is used (say, to compute ground motions). In such cases, published uncertainty bounds (e.g., M<sub>w</sub>6.8-M<sub>w</sub>7.0) are ranges of the estimated *median* magnitude considering one source of uncertainty. This may be distinctly different from the total uncertainty of the 1886 magnitude, which could be much greater. Consider, for example, that *instrumental* magnitudes have 95% CIs exceeding some of the ranges in Table 1 (e.g., Werner and Sornette, 2008).

222 In this study, and analogous to the analysis of MMI data, probabilistic site-adjusted ground motion IMs 223 will be predicted by 17 ENA GMMs. Conditioned on these IMs, the probability of liquefaction manifestation will be computed by the fragility functions of Geyin and Maurer (2020), wherein multiple 224 models for soil aging effects will be ensembled. These predictions will be repeated for a multitude of 225 226 hypothetical sources to compute the likelihood that each would produce the observed regional liquefaction response. Like the study of MMI data, this will result in a probability distribution of earthquake magnitude 227 228 and, arguably, a more complete and transparent understanding of the 1886 source model, at least insofar as 229 can be gained from the macroseismic evidence available for analysis.

### 230 3. Macrosesimic data

### 231 *3.1 MMI data*

Bakun et al. (2002) compiled 1,034 intensity reports from the 1886 event, including those of Dutton (1889) 232 233 and Bollinger and Stover (1976), and assigned MMI values per the: (i) MMI definitions of Wood and 234 Neumann (1931); and (ii) USGS National Earthquake Information Center practice for assigning intensity 235 (Stover and Coffman, 1993). In compiling these data, Bakun et al. (2002) either excluded MMI = 1 and MMI = 2 reports (i.e., documented instances where shaking was not felt) or found no such reports to 236 237 compile. Like all previous studies of these data, we assume that MMI values were accurately assigned by 238 Bakun et al. (2002), with proper consideration of structural vulnerability to avoid biasing estimates. While matters of MMI uncertainty and bias could be important and will be further discussed, they are largely 239 240 beyond the scope of the present analysis. It should be emphasized that our analysis is not an endorsement of the Bakun et al. (2002) dataset. Rather, we study this data because it is the most recent compilation of 241 242 1886 intensity reports and because many prior publications studied either the Bakun et al. (2002) 243 compilation, or earlier collections of data that Bakun et al. (2002) built upon. As previously stated, a future 244 study could further benefit from a complete reinterpretation of the more than 1000 original intensity reports. 245 While the coordinates of one MMI report from Ottawa, Ontario were judged to be erroneous and corrected by judgement, we otherwise adopt the Bakun et al. (2002) MMI data as presented therein. The locations of 246 these data are shown in Figure 3 and extend ~1500 km from Charleston, SC. 247

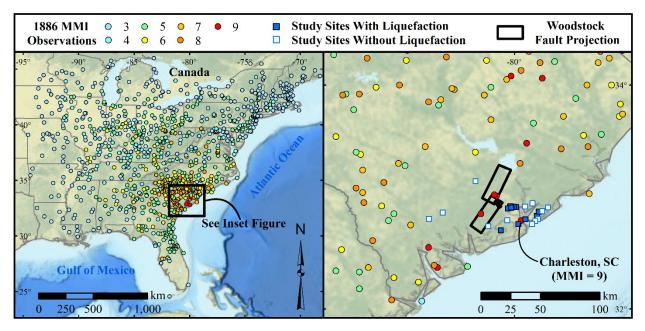


Figure 3. 1886 MMI observations (Bakun et al., 2002) and liquefaction evidence, as introduced subsequently. Also shown is the Woodstock Fault projection (Durá-Gómez and Talwani, 2009a,b).

248 3.2 Liquefaction data

Twenty-four sites where liquefaction manifestations were or were not observed in 1886, and where CPT testing was subsequently performed, will be studied. Liquefaction manifested (typically in the form of large sand boils) at twelve of these sites, whereas no evidence of liquefaction was observed at the remainder. These sites are summarized in Table 2, where citations are provided for the observed liquefaction response and for the geotechnical tests, which are all available in the public domain. These study sites are also mapped later in the paper.

CPT ID	Longitude	Latitude	Manifestation	Geotechnical Reference	Liquefaction Reference
BKY07	-79.9061	32.9150	No	USGS (2021)	Dutton (1889)
BKY09	-79.8385	32.9443	No	USGS (2021)	Dutton (1889)
BKY23	-79.9855	32.9115	Yes	USGS (2021)	Dutton (1889)
BKY24	-80.0071	32.9118	Yes	USGS (2021)	Dutton (1889)
CHN01	-79.7900	32.8030	No	USGS (2021)	Dutton (1889)
CHN07	-79.8134	32.7874	No	USGS (2021)	Dutton (1889)
CHN12	-79.7989	32.8310	Yes	USGS (2021)	Amick et al. (1990)
CHN15	-79.6998	32.9073	No	USGS (2021)	Dutton (1889)
CHN28	-79.8428	32.7597	No	USGS (2021)	Dutton (1889)
CHN29	-79.7840	32.8682	Yes	USGS (2021)	Amick et al. (1990)
CHN31	-79.7520	32.8741	No	USGS (2021)	Dutton (1889)

 Table 2. Summary of 1886 liquefaction data analyzed herein.

CHN32	-80.0267	32.9043	Yes	USGS (2021)	Dutton (1889)
CHN33	-80.0323	32.9166	Yes	USGS (2021)	Dutton (1889)
CHN34	-80.0391	32.9170	Yes	USGS (2021)	Dutton (1889)
CHN50	-80.1235	32.7023	Yes	USGS (2021)	Dutton (1889)
CHN59	-79.9655	32.7575	Yes	USGS (2021)	Dutton (1889)
CHN64	-80.0604	32.8982	Yes	USGS (2021)	Dutton (1889)
CREC1	-80.0655	32.7921	No	Boller (2008)	Martin and Clough (1994); Boller (2008)
FHS3	-80.3507	33.1420	No	Hasek (2016)	Williamson and Gassman (2014); Hasek (2016)
HA74	-80.0300	32.9050	Yes	Heidari & Andrus (2012)	Dutton (1889)
HWD2	-80.2355	32.7394	No	Hasek (2016)	Talwani and Cox (1985); Hasek (2016)
S99634DS1	-79.9015	32.8017	Yes	Heidari & Andrus (2010)	Dutton (1889)
WLC2	-80.6444	32.8607	No	Geiger (2010)	Geiger (2010)
WRAP2	-80.7666	32.8040	No	Geiger (2010)	Geiger (2010)

255 In compiling these study sites, an interpreted, potential lateral spread at Fort Dorchester, SC (Talwani 256 et al., 2011) was omitted because the liquefaction response was ambiguous and because lateral spreading – a distinct and complex manifestation of liquefaction – is not intended to be predicted by the liquefaction 257 258 model adopted herein. Whereas the MMI data are spread throughout ENA, the compiled liquefaction data 259 are within 100 km of Charleston. Because liquefaction is mechanistically implausible, irrespective of soil properties, at PGAs less than ~0.09 g (de Magistris, 2013), the maximum site-to-source distance of 260 261 liquefaction observations is inherently limited. By corollary, an analysis of sites where liquefaction was not 262 observed, but which are very far from Charleston, would not provide meaningful constraint of the source 263 model, given that the computed probability of such an observation is 100% even for very large earthquakes. 264 In addition, while the MMI data from 1886 are unlikely to grow significantly, additional liquefaction data 265 could be compiled. That is, the liquefaction response was documented in 1886 at more than twenty-four 266 sites, but costly geotechnical testing must also be performed at each site. In this regard, it is known that 267 CPTs have been performed near additional sites of observation, yet these data are privately held and could 268 not be obtained for analysis. Nonetheless, a larger dataset could be studied in the future.

## 269 **4. Methodology**

The methodology that will be used to analyze macroseismic data was first introduced by Rasanen and Maurer (2021, 2022), who collectively demonstrated and validated its use on landslide and liquefaction evidence produced by eleven modern earthquakes with known source models. This is the first application of the method to a prehistoric or pre-instrumental earthquake. The methodology will be covered in two sections. The first provides a succinct conceptual overview. The second describes in detail the applicationof the methodology to the 1886 Charleston earthquake.

#### 276 *4.1 Conceptual Overview*

Our goal is to probabilistically constrain the 1886 source model. This is accomplished by computing the likelihood that a rupture with some magnitude, location, and geometry, would produce a set of field observations (reported MMI or liquefaction observations), wherein uncertainties inherent to these observed outcomes are considered. In general, the likelihood of a parameter having some value, given a set of observations, is the product of the probabilities of those observations, conditioned on the parameter value. In other words, the likelihood of a rupture having some magnitude ( $M_w$ ), location (L), and geometry (G), given a set (x) of field observations at N different sites, can be computed as:

284 Likelihood 
$$(L, G, M_w | x) = P(X = x | L, G, M_w) = \prod_{i=1}^{N} P(X_i = x_i | L, G, M_w)$$
 (1)

where  $P(X_i = x_i | L, G, M_w)$  is the probability of what was observed at site *i* (i.e., the observed MMI or liquefaction response) given an earthquake with parameters *L*, *G*, and *M\_w*. By repeating for enumerable possibilities, the actual rupture parameters may be probabilistically constrained by the likelihood function (product of the probabilities of *N* observations), such that different source models will be found more and less likely to produce the observed evidence.

290 If the evidence is MMI data, then the probability of any one field observation (obs) is:

291 
$$P(\text{obs}|\text{EQK}: L, G, M_w) = \int_{IM} \int_{MMI} P(\text{obs}|MMI) f(MMI|IM) f(IM|L, G, M_w) \cdot dMMI \cdot dIM$$
(2)

292 where  $f(IM|L, G, M_w)$  is the probability of an IM value conditioned on fault parameters L, G and  $M_w$ , and 293 site parameter  $V_{S30}$  (i.e., the time averaged shear-wave velocity over the upper 30 m), as computed by a GMM that considers site response; f(MMI|IM) is the probability of an MMI value conditioned on the IM 294 value, and possibly on other parameters, as computed by an IM-MMI model; and P(obs|MMI) is the 295 296 binomial probability that the predicted MMI is equal to the observed MMI. In this work, we consider 297 predictions and observations to agree if within  $\pm$  0.5 MMI (e.g., predicted MMIs of 5.51 and 6.49 agree with an observed MMI of 6). Thus, we assign an implicit, uniform measurement uncertainty of  $\pm 0.5$  MMI, 298 299 but do not otherwise model the uncertainty of reported MMIs. In this regard, a thorough reinterpretation of the original intensity data, to include assignment of site-specific uncertainty distributions and corrections 300 301 for bias, would be a valuable endeavor. In the current effort, however, we adopt the Bakun et al. (2002) 302 MMI dataset, which has been studied by other modern investigators of the 1886 earthquake.

303 If the evidence is observed liquefaction response rather than MMI, then the probability of field 304 observation is computed per Eq (3) if liquefaction manifestation was observed and per Eq (4) otherwise:

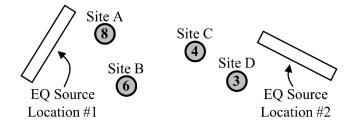
$$P(\text{Manifestation}|\text{EQK}: L, G, M_w) = \int_{IM} P(\text{Manifestation}|IM, M_w) f(IM|L, G, M_w) \cdot dIM$$
(3)

306 
$$P(\text{No Manifestation}|\text{EQK}: L, G, M_w) = 1 - \int_{IM} P(\text{Manifestation}|IM, M_w) f(IM|L, G, M_w) \cdot dIM$$
(4)

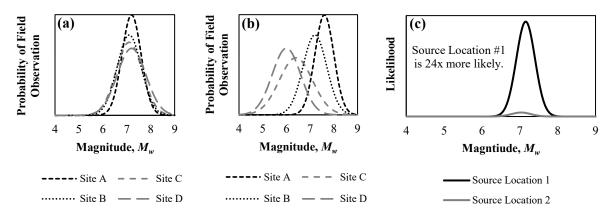
305

In these equations, uncertainty is not assigned to the observed liquefaction response;  $f(IM|L, G, M_w)$  has the same meaning as above; and  $P(\text{Manifestation}|IM, M_w)$  is computed by a probabilistic model that predicts the triggering of liquefaction at depth and its subsequent manifestation at the ground surface using subsurface geotechnical data. Ultimately, the uncertainties that are, and are not, accounted for in this work will be explicitly discussed.

312 Application of this method is demonstrated conceptually in Figure 4 considering four MMI observations and two hypothetical sources for the earthquake that produced the observations. In actual 313 314 analyses, a very large number of sources is considered. Figure 5 illustrates how the relative likelihoods of these two sources are assessed. Shown in Figure 5a are the computed probabilities of individual 315 observations, given a rupture of source one, as computed by Eq. 2 for varying  $M_w$ . In Figure 5b, this is 316 repeated considering a rupture of source two. In Figure 5c, the likelihood of each source is computed as a 317 318 function of  $M_w$  by Eq. 1 (i.e., the product of the four probability distributions in Figure 5a or 5b). In this simple example, source one has a far greater peak likelihood of producing the set of field observations, 319 320 whereas source two is very unlikely to do so, regardless of its  $M_{w}$ . By repeating this process for an array of 321 hypothetical sources and generating a likelihood distribution for each, the characteristics of the causative 322 rupture (e.g., location, orientation, magnitude) are probabilistically constrained, to the degree that evidence 323 permits. In this work, we compute  $M_{\psi}$  probability-distributions conditioned on both the hypothesized Woodstock fault and on an unknown source. The latter is accomplished by aggregating probability 324 distributions from all hypothetical sources and thus includes the uncertainties of fault location and 325 geometry, whereas the former assumes the fault location and geometry are known. 326



**Figure 4.** Hypothetical MMI inverse analysis consisting of four field sites, where MMI values were reported for each of the site locations (the value in the circle). In addition, two hypothetical sources for the earthquake that produced these observations are shown.



**Figure 5.** Approach for computing the likelihood of the sources depicted in Figure 4: (a) probabilities of individual observations, given an earthquake at location one of variable  $M_w$ ; (b) probabilities of individual observations, given an earthquake at location two of variable  $M_w$ ; (c)  $M_w$ -likelihood distributions for source locations one and two.

### 327 4.2 Implementation Details

The implementation of Eqs. 1-4 is next described in detail. This includes both general methods transferrable to other regions and to other forms of evidence, as well as the specific models adopted for the 1886 earthquake. Because the analyses of MMI and liquefaction data are procedurally similar, we first fully describe the former and then succinctly discuss differences specific to the latter. For completeness, however, the liquefaction analysis procedure is fully replicated in the electronic supplement.

As detailed in the following, two approaches will be used to model hypothetical seismic sources. These 333 334 sources are first treated as earthquake epicenters on a grid pattern, which we term the "epicenter search". To use modern GMMs, which are applicable only to faults, each epicenter is converted to an amorphous 335 336 fault realization using site-to-source distance correlations, which implicitly assume some fault geometry as 337 a function of magnitude. With this approach, the source location may be investigated (i.e., probabilistically constrained) without consideration of rupture geometry. While this reduces computational expense, it 338 considers only a "median" geometry and thus omits one source of uncertainty. Accordingly, and following 339 340 constraint of the epicentral region, seismic sources are next treated as faults having an array of locations, 341 lengths, and orientations, which we term the "fault search." Of these enumerable faults, one is the 342 hypothesized Woodstock Fault, which we highlight in the results for obvious reasons.

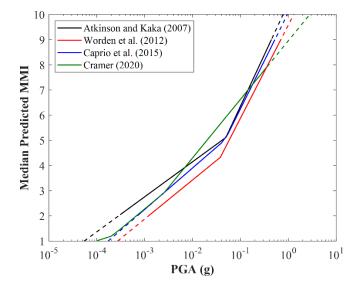
### 343 4.2.1 MMI Analysis Procedure

344 The analysis of MMI data is completed via the following 15 steps:

- 345 (1) Create an array of seismic sources. In this paper, a 62,500 km<sup>2</sup> grid of surficial points was centered on
- 346 Charleston. Within this grid, a finer point spacing increases spatial resolution while a coarser spacing

- decreases computational demand (which can be significant, and tractable only with high performance
- computing). Having found that 1 km and 10 km spacing produce nearly identical outcomes, we adopted
- 349 the latter spacing when studying MMI observations. These surficial points will be treated first as
- earthquake epicenters (i.e., the "epicenter search") and second as surface projections of the centroids of
- faults (i.e., the "fault search").
- 352 (2) Select N study sites (i.e., the MMI observations compiled by Bakun et al. (2002)).
- 353 (3) Select an appropriate GMM. In this paper, the 17 GMMs developed by the NGA East Project (Goulet et
- al., 2018) were coalesced in a logic tree using the weights proposed by Goulet et al. (2018). Using this
  scheme, the least and greatest model weights were ~2% and 10%, respectively.
- **356** (4) For each seismic source created in (1):
- 357 (5) For each seismic-source  $M_w$  considered (a range of M<sub>w</sub>4 to M<sub>w</sub>8.2 the applicable range of the adopted 358 GMMs – was used in this paper):
- 359 (6) For each of N study sites selected in (2), cycling from i = 1 to N:
- (7) Compute the site-to-source distance(s) required by the GMM chosen in (3), as measured from study site 360 361 *i* to the seismic source selected in (4). For the GMMs adopted herein, the only such metric required is-the 362 closest distance to fault rupture ( $R_{RUP}$ ). When the seismic sources in (1) are treated as epicenters, rather 363 than faults, the correlations of Scherbaum et al. (2004) were used to estimate a median  $R_{RUP}$  from 364 epicentral distance ( $R_{EPI}$ ). In effect, these correlations, which are magnitude dependent, convert each 365 epicenter from (1) into a median realization of a multidimensional fault. Alternative approaches to 366 estimating  $R_{RUP}$  from point sources are provided by Bommer et al. (2016) and Thompson and Worden (2018). Ultimately, the sources in (1) are explicitly modeled as faults to determine whether the field 367 368 evidence can constrain the 1886 source beyond a point location. In doing so,  $R_{RUP}$  is directly measured from fault planes and the uncertainty of fault orientation is considered. 369
- 370 (8) Using the GMM in (3),  $M_w$  from (5), and site-to-source distances from (7), compute the probability 371 density function (PDF) of expected *PGA* at site *i*, modified for site effects. In general, this PDF is a 372 lognormal random variable described by a median and lognormal standard deviation, which are given by
- a GMM. PGA predictions beyond  $\pm 3$  standard deviations of the median were truncated, as is typical, and
- the PDF was scaled such that the area beneath it was one. In this study, the ENA weighted GMM predicts
- 375 PGA for reference rock conditions. Accordingly, the  $V_{S30}$ -dependent model of Harmon et al. (2019),
- 376 which is ENA-specific and developed as part of the NGA-East Project, was used to adjust *PGA*s for local
- site effects.  $V_{S30}$  was estimated at the site of each field observation using the Heath et al. (2020) maps.
- 378 (9) For each possible *PGA* value at study site *i*, as computed in (8) for a given  $M_w$  and  $R_{RUP}$  pair:
- 379 (10) Select an IM-MMI model. In this paper, the Atkinson and Kaka (2007), Worden et al. (2012), Caprio et
- al. (2015), and Cramer (2020) models were adopted. The inputs to these models vary but generally include

381 PGA,  $M_w$ , and  $R_{RUP}$ . The median MMI predicted by each model is plotted in Figure 6 as a function of 382 PGA considering an M<sub>w</sub>6 event at  $R_{RUP} = 100$  km.



**Figure 6.** Predicted median MMI versus PGA according to four recent models (where applicable,  $M_w = 6.0$ ;  $R_{RUP} = 100$  km). Solid lines indicate the range of each model's proposed applicability (typically the range of training data); dashed lines indicate extrapolation beyond these respective bounds.

- 383 (11) Using the IM-MMI model selected in (10), the  $R_{RUP}$  from (7), and the PGA from (9), compute the PDF
- of expected *MMI* at study site *i*. In this paper, *MMI* predictions beyond  $1 \le MMI \le 10$  were truncated

because the definition of *MMI* gives a lower bound of 1 and because intensities greater than 10 are rarely

assigned in practice (Stover and Coffman, 1993). Following truncation (when applicable) the PDF was

- scaled such that the area underneath it was one.
- 388 (12) For each possible MMI value at study site *i*, as computed in (11):
- 389 (13) Compute the probability of field observation as described in Eq. 2. Completing this equation (i.e., by
- multiplying the probability of field observation by the probabilities of *MMI* and *PGA*, then summing over
- all *MMI* and *PGA* values) gives the probability of the field observation at site *i* for a given seismic source
- location, geometry, and  $M_w$ . Repeating steps 6-13 for all  $M_w$  results in a probability of field observation
- curve for each source, examples of which are in Figures 5a and 5b.
- 394 (14) Compute the likelihood of a seismic source (as evidenced by MMI) as a function of  $M_w$  by multiplying
- the probabilities of all field observations (i.e., multiply the curves in Figures 5a or 5b at each value of
- 396  $M_w$ ). The result, an example of which is shown in Figure 5c, is a likelihood distribution of  $M_w$  conditioned
- 397 on a single seismic source (i.e., assuming that a fault with given location and geometry, but unknown  $M_{w}$ ,
- is responsible for producing the field observations).

399 (15) Repeating steps 5-14 for all seismic sources created in (1) results in a likelihood distribution of  $M_{\rm W}$  for 400 each. Collectively, this field of distributions describes the locations and magnitudes of earthquakes that 401 are, and are not, likely to produce the field evidence. To allow for relative likelihoods to be compared visually, we normalize the likelihood of each source by the peak likelihood among all sources, such that 402 403 the most likely source has a normalized peak value of one. We then map contours of likelihood to identify this location. The  $M_{\rm W}$  distribution at this location (see Figure 5c) is the PDF of the inverted  $M_{\rm W}$  conditioned 404 405 on the most likely source. Finally, by aggregating PDFs from all potential sources in (1), an overall PDF of  $M_{\nu}$ , considering all possible sources, is produced. While a single source will always be "most likely," 406 407 earthquakes at other locations typically also have potential to produce the evidence. This latter PDF, 408 conditioned on all possible sources, includes that uncertainty.

## 409 *4.2.2 Liquefaction Analysis Procedure*

The analysis of liquefaction is akin to that of MMI, differing only in steps (10) through (13) of the 410 411 preceding. Once site-adjusted PGAs are predicted at sites of observation (in this case, sites with liquefaction 412 observations), the factor of safety against liquefaction is computed throughout the soil profile using the 413 Boulanger and Idriss (2014) CPT-based triggering model, which is a function of PGA,  $M_{w}$ , and subsurface 414 geotechnical data. As part of this procedure, the effects of soil aging are separately accounted for with three proposed models, in addition to a control analysis without any such accounting. Two of these models, 415 416 henceforth termed "A" and "B", are regional-scale corrections based on the Weems et al. (2014) geology map of the Charleston region, whereas Model "C" uses site-specific measurements. Model A uses the 417 418 Hayati and Andrus (2009) aging-correction model based on the measured to estimated shear-wave velocity 419 ratio (MEVR). As MEVR increases, the shear stiffness of soil measured at small strain exceeds that inferred 420 from large strain measurements, which may be interpreted as an indicator of cementation. Median MEVR 421 values for each geologic unit were adopted from the regional sampling of Heidari and Andrus (2012). Model 422 B uses the Hayati and Andrus (2009) correction model based on depositional age, which we estimated for each geologic unit from the Weems et al. (2014) map. As a deposit's age increases, an increasingly larger 423 correction is applied to the computed liquefaction resistance. Lastly, Model "C" again adopts the MEVR-424 425 based approach of Hayati and Andrus (2009), but in this case MEVR is directly computed at each study site using data from seismic CPTs (i.e., CPT resistances and shear-wave travel times) per the method of 426 427 Andrus et al. (2009). This approach produced corrections that tended to exceed those of Models A and B. 428 While corrections from Model C – being based on site-specific subsurface data – should be most efficient, 429 all such corrections are likely very uncertain and the best approach to account for soil aging is actively 430 debated. We thus argue that all three models, which scale upward the computed liquefaction triggering 431 resistance at depth, warrant consideration. However, given that the field observations are of surface

manifestations (i.e., liquefaction vents and dikes) rather than observations of liquefaction at discrete depths,
surface manifestations must be predicted for proper comparison to field observations. Accordingly, the
liquefaction potential index (*LPI*) of Iwasaki et al. (1978) was adopted, given its longstanding use:

435 
$$LPI = \int_0^{20 m} F(FS_{liq}) \times w(z)dz$$
(5)

where  $F(FS_{liq})$  and w(z) weight the respective influences of  $FS_{liq}$  and depth, z, on surface manifestation. 436 437 In brief, LPI assumes that manifestation depends on the FS<sub>liq</sub> in each soil stratum, the thickness of all liquefied strata, and the proximity of those strata to the surface. Given this definition, LPI can range from 438 0 to 100. Next, the probability of liquefaction manifestation was computed at each study site, considering 439 all values of  $M_w$  from (5) and all values of PGA from (9), using the fragility function of Geyin and Maurer 440 (2020), which relates the probability of liquefaction manifestation to LPI, and which was trained on all 441 442 globally available CPT-based liquefaction case histories. Lastly, and analogous to (13), the probability of what was observed in the field was computed by Eq. 3 or 4, depending on whether manifestations were or 443 444 were not observed. The liquefaction analysis procedure is fully replicated in the electronic supplement.

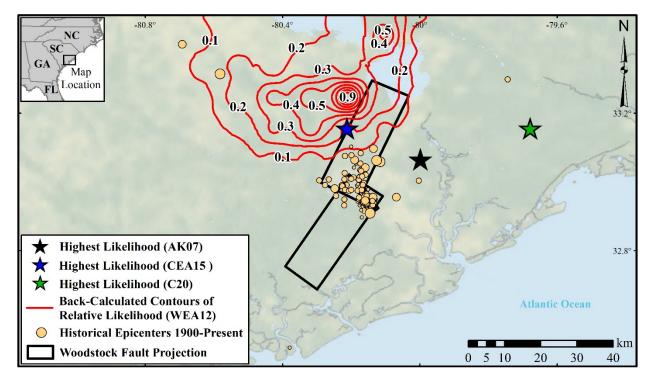
### 445 5. Results and discussion

Using the above data and methodologies, the source of the 1886 Charleston rupture was rigorously 446 447 investigated. Results are separately presented for MMI and liquefaction evidence. For each, three types of analyses were performed, namely the: (i) "epicenter search", wherein seismic sources were modeled as 448 points; (ii) "fault search," wherein seismic sources were modeled as faults; and (iii) "Woodstock Fault," 449 450 wherein the hypothesized Woodstock Fault is directly considered. The first provides preliminary probability 451 distributions of the source location and magnitude. The second provides final distributions that include 452 additional source-model uncertainties and conveys whether the orientation of the causative fault can be constrained. The third provides results conditioned on the singular Woodstock Fault and should be adopted 453 454 if all other seismic sources, known and unknown, are dismissed as sources of the 1886 earthquake.

### 455 *5.1 Analysis of MMI Data: Epicenter Search (Seismic Sources Modelled as Points)*

Treating seismic sources as points, the inversion methodology was applied using the Atkinson and Kaka (2007), Worden et al. (2012), Caprio et al. (2015), and Cramer (2020) IM-MMI models. For brevity, these models are henceforth titled AK07, WEA12, CEA15, and C20. For these analyses, MMI observations were studied if within 600 km of Charleston. While the sensitivity of this decision will be analyzed further, it was made based on a trait of the Bakun et al. (2002) MMI dataset that we view as problematic, but which prior researchers have not discussed. Namely, Bakun et al. (2002) did not include "negative" observations (i.e., MMI values of 1 or 2) where ground motions were not perceived. This is notable, given that a 463 *distribution* of MMI values is naturally experienced at any given distance. Problematically, the omission of 464 low MMI values asymmetrically truncates this distribution at large distances (i.e., those where some 465 observers report MMI 1 or 2). The analysis of this dataset at large distances thus potentially has the effect of biasing inverted magnitudes toward larger values, since those magnitudes need not adhere to small MMI 466 values that were experienced, but which were systematically undocumented. The analysis of data only at 467 small distances is equally undesirable, of course, given the inherent benefits of analyzing more data over a 468 469 wider spectrum of attenuation. Based on these competing interests and given that we begin to observe 470 evidence of MMI truncation at distances exceeding 600 km, this threshold was provisionally selected.

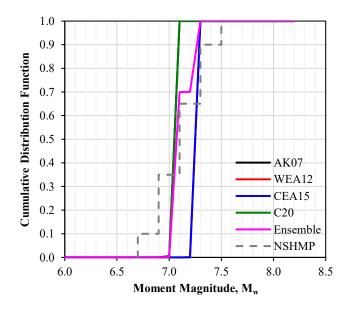
The resulting geospatial contours of seismic-source likelihood are shown in Figure 7 for the WEA12 471 472 model. Also shown is the surface projection of the hypothesized Woodstock fault (Durá-Gómez and 473 Talwani, 2009a,b), as well as historical epicenters ( $M_w > 1.0$ ; 1900-present). These contours show the source locations most and least likely to produce the MMI evidence. Using the WEA12 model, a zone of 474 high likelihood is computed just northwest of the Woodstock Fault projection, with the most likely source 475 476 located 2 km from the mapped projection. By contrast, source locations to the south and east are 477 comparatively very unlikely to produce the MMI observations. For brevity, results using the AK07, CEA15, and C20 models are summarized in Figure 7 via stars, which denote the epicenters deemed most likely by 478 479 each model. Complete contour maps, analogous to those in Figure 7 for WEA12, are provided in the 480 electronic supplement. While CEA15 produces a result similar to WEA12, with the most likely source 481 inside the Woodstock Fault projection, AK07 and C20 suggest most likely sources ~10 km and ~35 km 482 east of the projection. Thus, all analyses point to a source north of Charleston. Some IM-MMI models strongly support an earthquake source in the immediate vicinity of the hypothesized Woodstock Fault, while 483 484 others do not.



**Figure 7.** Likelihood contours produced using the WEA12 model which indicate source locations that are, and are not, likely of producing the MMI evidence. Stars = the most likely source locations per the AK07, CEA15, and C20 models; AK07 = Atkinson and Kaka (2007), WEA12 = Worden et al. (2012), CEA15 = Caprio et al. (2015), C20 = Cramer (2020); black rectangles = the Woodstock fault projection (Durá-Gómez and Talwani 2009a,b); tan circles = historical epicenters ( $M_w > 1.0$ ; 1900-present) (USGS, 2022).

485 Shown in Figure 8 are the cumulative distribution functions (CDFs) of M<sub>w</sub> inverted from MMI data, as 486 computed with each of the four IM-MMI models. These results include the uncertainty of an unknown 487 source (epicenter) location, the uncertainties of expected shaking intensities conditioned on a given source, considering also the epistemic uncertainty of GMM selection, and the uncertainties of the IM-MMI models. 488 For comparison, the CDF assigned to the Charleston source in the NSHMP (Petersen et al., 2014, 2020) is 489 490 also shown and ranges from M<sub>w</sub>6.7 to M<sub>w</sub>7.5 with a median of M<sub>w</sub>7.1. AK07 and C20 produce nearly identical CDFs with a median of M<sub>w</sub>7.1. WEA12 and CEA15 produce similarly identical CDFs, but with a 491 492 median of  $M_w$ 7.3. The grouping of these outcomes can be traced to Figure 6. AK07 and C20 tend to predict larger MMIs for a given PGA. As a result, smaller magnitude ruptures result from the use of AK07 and 493 494 C20 within the inversion methodology. Each of the individual CDFs has relatively low uncertainty. This 495 can be attributed to: (i) the large quantity of MMI observations; (ii) the fact that not all uncertainties are considered (e.g., those of site conditions or rupture geometry, among others); and (iii) the relatively large 496 variance of MMI observations over any given site-to-source distance. Regarding the last, extreme outliers 497 498 from the mean MMI trend have the effect of yielding very low likelihoods for small and large magnitude

499 events. That is, only a narrow range of magnitudes are likely to simultaneously produce MMI observations 500 of, say, 3, 5, and 7 at the same site-to-source distance and on similar site profiles. Conversely, a wider range 501 of magnitudes could simultaneously produce MMI observations of 4, 5, and 6. While geostatistical analyses could conceivably identify and delete observations that are extreme outliers, the justification would be 502 largely speculative without an intensive reinvestigation. In other words, we are unaware of any objective 503 basis for deleting some MMI reports but not others in the absence of a complete reinterpretation of the more 504 505 than 1000 original intensity reports. Lastly, to include the epistemic uncertainty of IM-MMI model 506 selection, each CDF was weighted to produce an ensemble CDF, as shown in Figure 8. This selection of 507 weights is heuristic, as no quantitative justification could be identified (e.g., based on model residuals in 508 ENA). The weights and justifications are as follows. C20 (0.4) is the latest ENA-specific model and was 509 trained using the largest quantity of ENA data. AK07 (0.3) is also ENA-specific, but also found no need 510 for region-specific models within North America. For that reason, and because differences in methodology (e.g., data selection, intensity scales, regression techniques) can result in very different models with 511 512 apparently similar performance (i.e., on the respective training sets), we hesitate to omit other reputable 513 models even if nonspecific to ENA. Accordingly, WEA12 (0.15), which was trained only on California 514 data, and CEA15 (0.15), which was trained on global data including some from ENA, are also weighted in 515 the ensemble. While CEA15 might otherwise warrant greater weighting, its training set was truncated at a 516 site-to-source distance of 200 km, which calls into question its suitability in ENA, where ground motions 517 are commonly felt at greater distances (as in the 1886 event).



**Figure 8.** CDFs of rupture magnitude, as inverted from MMI observations using four IM-MMI models, wherein seismic sources are modelled as points. AK07 = Atkinson and Kaka (2007), WEA12 = Worden et al. (2012), CEA15 = Caprio et al. (2015), C20 = Cramer (2020), and NSHMP = National Seismic Hazard Model Project (Petersen et al., 2014, 2020).

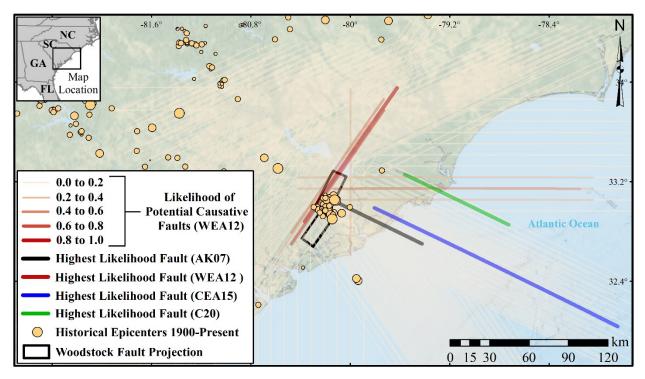
#### 518 5.2 Analysis of MMI Data: Fault Search (Seismic Sources Modelled as Faults)

519 To determine whether MMI observations can constrain the seismic source beyond a point location, the point 520 sources were next treated as surface projections of the centroids of faults. For simplicity, the faults were initially assumed to be strike-slip with a dip of 90 degrees. The orientations of the faults were discretized 521 into strike azimuths of  $30^{\circ}$  increments (i.e.,  $0^{\circ}$ ,  $30^{\circ}$ ,  $60^{\circ}$ ...). The lengths of the faults were computed via the 522 magnitude-dependent correlation of Wells and Coppersmith (1994). The depth to the top of rupture was 523 524 estimated using the correlation of Kaklamanos et al. (2011), which produces estimates consistent with the 525 inferred depths of modern ruptures in the region (Chapman et al., 2016). Thus, while not every aspect of rupture geometry was assigned uncertainty, the analyses can nonetheless determine whether some fault 526 527 alignments (e.g., the hypothesized Woodstock Fault) are more likely than others to produce the observed 528 evidence. With this approach, the inversion methodology was again applied to the Bakun et al. (2002) 529 dataset using each of the four IM-MMI models.

Results are shown in Figure 9 for the WEA12 model in the style of a heat map. Faults more likely to 530 531 produce the MMI observations have thicker lineweight and are darker in color. Faults very unlikely to 532 produce the evidence have thin lineweight and light color, and thus blend with the map's background. Also shown is the surface projection of the hypothesized Woodstock Fault (Durá-Gómez and Talwani 2009a,b) 533 534 as well as historical epicenters. While most modelled faults are relatively unlikely to produce the evidence, 535 the faults deemed most likely align with the hypothesized Woodstock Fault, albeit their magnitudes are 536 greater than most prior estimates. The singular fault most likely to produce the MMI evidence, for example, 537 has a median M<sub>w</sub> of 7.60. Faults striking E-W and N-S have lesser, but nontrivial, likelihoods of producing 538 the MMI evidence.

539 For brevity, results using the AK07, CEA15, and C20 models are each summarized in Figure 9 by single lines, which denote the singular faults deemed most likely by each model. Complete heat maps, 540 analogous to those in Figure 9 for WEA12, are provided in the electronic supplement. These three models 541 produce results that are similar to one another and different from WEA12, with the most likely faults 542 oriented perpendicular to the Woodstock Fault and located partly offshore. These most-likely faults range 543 in median magnitude from M<sub>w</sub>7.20 to M<sub>w</sub>7.90. A close inspection of these results (electronic supplement) 544 reveals that each of these models finds faults consistent with the Woodstock Fault to have low relative 545 546 likelihoods of producing the MMI evidence, as compared to all other hypothetical faults. It is worth noting 547 that the two ENA-specific IM-MMI models (AK07 and C20) point to faults inconsistent with the 548 Woodstock Fault. While these results do not necessarily indicate that the Woodstock Fault was not 549 responsible for the 1886 earthquake, they do indicate that the causative fault cannot currently be constrained 550 by the available MMI evidence. That is, faults striking SE and partly offshore are just as likely to produce 551 this evidence as faults striking NE and onshore. While different assumptions about fault depth, length, and

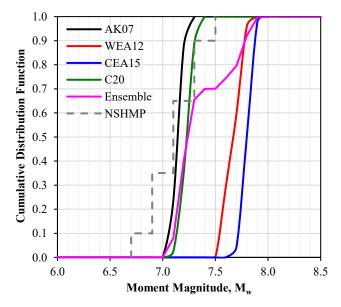
552 dip invariably change the results, a parametric analysis indicates that these changes are minor, upholding 553 the overall conclusion that the MMI data cannot constrain the 1886 fault rupture. If anything, the data point 554 to a fault striking SE and partly offshore. This conclusion might be different if: (i) more MMI observations were available in the near field; and (ii) the affected area was not on a coastline. Analogous to the inversion 555 of an epicenter from wave-arrival times, which relies on distributed instruments, this inversion relies on 556 distributed observers to "record" the amplitudes of ground motions. Thus, in events such as this, where the 557 558 near-field is not more densely populated with observations than the far field, and where observations are 559 geographically asymmetric, it may be more difficult to detect directions and rates of ground-motion 560 attenuation, and thus more difficult to constrain the causative fault from which motions propagated.



**Figure 9.** Likelihoods of hypothetical faults with differing orientations and lengths to produce the MMI evidence, as obtained using the WEA12 model. Also shown are the singular faults deemed most likely using the AK07, CEA15, and C20 models; AK07 = Atkinson and Kaka (2007), WEA12 = Worden et al. (2012), CEA15 = Caprio et al. (2015), C20 = Cramer (2020); black rectangles = the Woodstock projection (Durá-Gómez and Talwani 2009a,b); tan circles = historical epicenters ( $M_w$ >1.0; 1900-present) (USGS, 2022).

561 Shown in Figure 10 are the CDFs of M<sub>w</sub> inverted from MMI data via the "fault search", as computed 562 with each of the four IM-MMI models. These results include the uncertainty of an unknown source location, 563 the uncertainty of unknown fault orientation, the uncertainties of expected shaking intensities conditioned 564 on a given source and considering the epistemic uncertainty of GMM selection, and the uncertainties of the 565 IM-MMI models. It should be emphasized that these results do not include every source of uncertainty. 566 Omitted, for example, are the uncertainties of: (i) fault depth and length; (ii) fault dip; and (iii) site  $V_{S30}$ . For 567 each of these inputs only a median prediction was used. For comparison, the CDF assigned to the Charleston source in the NSHMP (Petersen et al., 2014, 2020) is also shown. Like the "epicenter search" results in 568 Figure 8, the AK07 and C20 models produce similar CDFs with a lesser median of  $\sim M_w 7.2$ , while WEA12 569 and CEA15 produce similar CDFs with a greater median of  $\sim M_w 7.7$ . It is readily apparent that the epistemic 570 571 uncertainty of IM-MMI model selection is considerable. To account for this uncertainty, each CDF was 572 weighted per the prior scheme to produce an ensemble CDF, as shown in Figure 10, which has a median of  $\sim$ M<sub>w</sub>7.25 and 95% CI of M<sub>w</sub>7.05 to M<sub>w</sub>7.85. In this regard, efforts to better quantify the suitability of various 573 574 IM-MMI models to ENA, and in turn to refine the weighting scheme used herein, could have significant 575 influence on the overall conclusions of this study (and presumably also on studies of other seismic sources). 576 Despite this ambiguity, the ensemble CDF in Figure 10 suggests that ruptures larger than  $M_w7.5$  – the largest value considered in the NSHMP - could produce the observed evidence. 577

Of course, it should also be emphasized that the numerous hypothetical faults considered in our methodology and aggregated to form the results in Figure 10 may not exist. That is, some of these gridded faults are more likely to produce the 1886 MMI data than the hypothesized Woodstock Fault, but these various faults are not necessarily present. Conversely, there is evidence (e.g., seismological data, geophysical surveys) supporting the Woodstock Fault's existence (Durá-Gómez and Talwani 2009a,b; Chapman et al., 2016). Accordingly, the preceding analyses are next repeated, conditioned solely on the Woodstock Fault, thereby removing all source uncertainties aside from magnitude.



**Figure 10.** CDFs of rupture magnitude, as inverted from MMI observations using four IM-MMI models, wherein seismic sources are modelled as faults. AK07 = Atkinson and Kaka (2007), WEA12 = Worden et

al. (2012), CEA15 = Caprio et al. (2015), C20 = Cramer (2020), and NSHMP = National Seismic Hazard Model Project (Petersen et al., 2014, 2020).

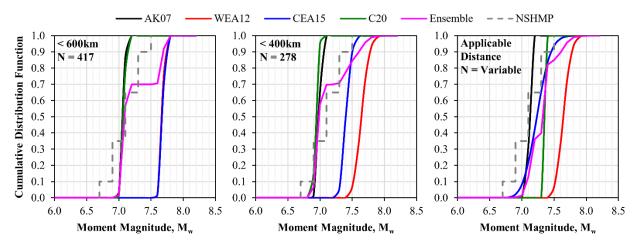
### 585 5.3 Analysis of MMI Data: Woodstock Fault

586 While the parameters of the Woodstock Fault are uncertain, we adopt the source model of Durá-Gómez and 587 Talwani (2009a,b) exactly as proposed therein, the surface projection of which appears in Figures 1, 7, and 9. While Chapman et al. (2016) do not identify the Woodstock Fault by name, they infer and describe from 588 589 more recent seismological records a source that is very similar: "We interpret ... that the modern seismicity 590 is the lingering aftershock sequence of the 1886 shock and that the mainshock occurred on a south-striking, 591 west-dipping fault plane ..." To evaluate the prior site-to-source distance threshold of 600 km (i.e., that 592 which was used to exclude MMI observations), these analyses were also repeated using MMI data at 593 different site-to-source distances. Shown in Figure 11, for example, are the CDFs of  $M_w$  conditioned on the 594 Woodstock Fault and inverted from MMI data in two distance bins: (a)  $\leq 600$  km; and (b)  $\leq 400$  km. Several 595 observations are made from Figure 11 as follows.

First, it is observed that the ensemble CDF in Figure 11a is only marginally less uncertain than the 596 597 ensemble CDF in Figure 10, meaning that conditioning the analysis on the Woodstock Fault does not 598 significantly alter the results. In other words, the uncertainties of fault location and geometry are relatively 599 minor, given the data available for analysis. As previously discussed in Figure 9, these data are unable to 600 constrain the causative fault. That is, faults with diverse locations and orientations have similar likelihoods 601 of producing the MMI evidence, which may be due to a paucity of near-field observations. However, 602 because similar rupture magnitudes are inferred for these various faults irrespective of their positions, the CDFs conditioned on the Woodstock Fault and the CDFs conditioned on an unknown fault are similar. 603

Second, a site-to-source distance bias is observed from Figures 11a and 11b. While the CDFs become 604 605 more vertical as more observations are included (reflecting a decrease in uncertainty) the CDFs also tend 606 to increase in magnitude. Observations at distances up to 600 km suggest larger rupture magnitudes than 607 those at distances up to 400 km. The median magnitude, for example, increases by ~0.25Mw according to 608 CEA15, by ~0.1M<sub>w</sub> according AK07 and C20, and by ~0.01M<sub>w</sub> according to WEA12. It could be shown that the prior "epicenter search" and "fault search" results have a similar degree of sensitivity. This distance 609 610 bias could be present in either: (i) one or more of the component models utilized (e.g., the GMMs, site-611 response model, or IM-MMI models); or (ii) in the MMI data itself. With respect to the adopted models, all 612 were shown to be unbiased during their respective trainings and cannot be further tested in the absence of 613 additional data. It is worth noting again, however, that the CEA15 model, which exhibits the greatest siteto-source distance sensitivity in Figure 11, is herein applied to data much more distant than it was trained 614 615 on. It was for this reason that CEA15 was given low weighting despite having a large global training set. 616 With respect to the MMI data, we reiterate that analyzing more data over a wider spectrum of attenuation

617 is beneficial but remain adamant that the Bakun et al. (2002) dataset is apt to introduce bias at large 618 distances, given that small MMI values were systematically undocumented. To further probe the issue of 619 IM-MMI model applicability, the analyses were performed using only the data to which each model is applicable (as stated by the original authors, or otherwise interpreted by the present authors). Specifically, 620 the WEA12, AK07, CEA15, and C20 models were respectively applied to observations within site-to-621 source distances of 400 km, 800 km, 200 km, and 1500 km. The results are shown in Figure 11c. The most 622 623 salient changes are: (i) the CEA15 CDF is more uncertain (due to fewer field observations) and has a 624 reduced median of  $\sim M_w 7.2$ ; and (ii) the C20 CDF is less uncertain (due to more field observations) and has an increased median of ~Mw7.35. In summary, we are unsure why the results show evidence of site-to-625 source distance bias at distances less than 600 km. However, we argue that the bias observed at larger 626 627 distances (e.g., the C20 result in Figure 11c) is at least partly due to the aforementioned bias of the Bakun 628 et al. (2002) dataset. As such, we prefer not to glean new conclusions from Figure 11c or champion its 629 results over those in Figure 11a. Our preferred CDF conditioned on the Woodstock Fault thus has a median 630 of M<sub>w</sub>7.10 and 95% CI of M<sub>w</sub>7.0 to M<sub>w</sub>7.75. Given an unknown source, our preferred CDF (Figure 10) has median of ~Mw7.25 and 95% CI of Mw7.05 to Mw7.85. For either result, the epistemic uncertainty of IM-631 632 MMI model selection has significant influence.



**Figure 11.** CDFs of rupture magnitude, as inverted from MMI observations using four IM-MMI models and assuming the Woodstock Fault (Durá-Gómez and Talwani, 2009a,b) to be the earthquake source. AK07 = Atkinson and Kaka (2007), WEA12 = Worden et al. (2012), CEA15 = Caprio et al. (2015), C20 = Cramer (2020); NSHMP = National Seismic Hazard Model Project (Petersen et al., 2014, 2020).

#### 633 5.4 Analysis of Liquefaction Data: Epicenter Search (Seismic Sources Modelled as Points)

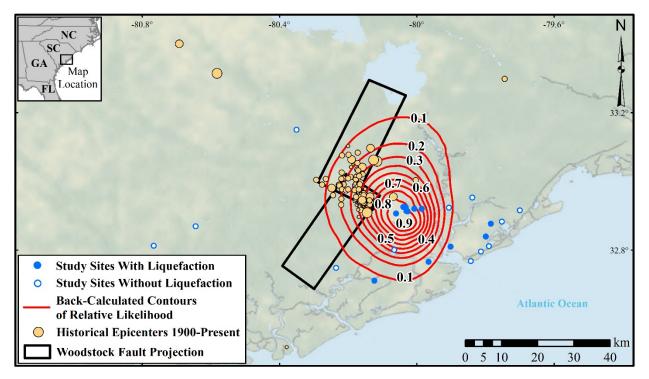
The inversion methodology was next applied to the 24 study sites summarized in Table 2 where the 1886

635 liquefaction response was observed, or has since been investigated, and where CPT testing has been

636 performed. Analogous to the analysis of MMI data, the seismic source was first treated as an epicenter with

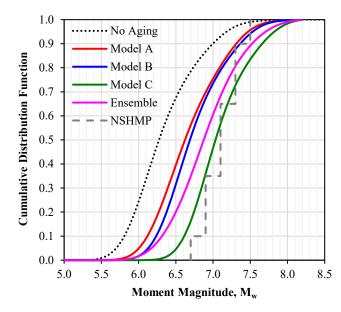
637 unknown location. The resulting contours of seismic-source likelihood, which depict the epicenters most 638 and least likely to produce the observed liquefaction response, are mapped in Figure 12, as are the 24 study 639 sites. Also shown are historical epicenters and the hypothesized Woodstock Fault (Durá-Gómez and Talwani, 2009a,b). These initial results include no correction to account for the effects of soil aging on 640 liquefaction. As shown in Figure 12, the analysis identified a source area east of the Woodstock fault. While 641 the single most likely epicenter is 9 km from the proposed fault projection, epicenters within the projection 642 have up to 70% relative likelihood. Notably, these geolocation results are largely insensitive to the treatment 643 644 of soil aging. Because aging correction models tend to adjust the computed liquefaction resistances by similar amounts, the inverted magnitude tends to be significantly affected, whereas the inverted, most likely 645 source location does not. For this reason, geolocation results are not shown for each aging-correction model. 646 647 All give similar results to those in Figure 12, indicating that an epicenter more than  $\sim 20$  km in any direction 648 from the northern Charleston Peninsula (e.g., the Charleston Airport) is relatively unlikely to produce the

649 observed evidence.



**Figure 12.** Likelihood contours indicating source locations that are, and are not, likely of producing the observed liquefaction response. Black rectangles = the Woodstock fault projection (Durá-Gómez and Talwani 2009a,b); tan circles = historical epicenters ( $M_w$ >1.0; 1900-present) (USGS, 2022).

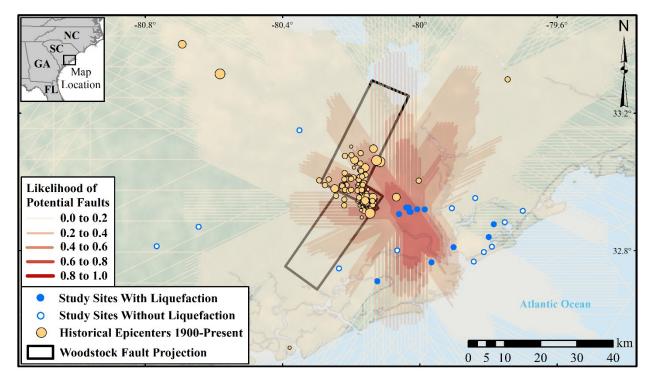
650 Shown in Figure 13 are the CDFs of  $M_w$  inverted from liquefaction evidence, as computed using: (i) no 651 aging correction; and (ii) each of the three correction methods discussed in the *Liquefaction Analysis*  652 Procedure (i.e., Models "A, B, and C"). It is apparent from Figure 13 that the uncertainties of whether and 653 how to correct for the effects of soil aging have significant influence, with the inverted median magnitudes 654 ranging from  $\sim M_w 6.3$  (no aging correction) to  $\sim M_w 7.0$  (Model C). Irrespective of soil aging, it is also apparent that a magnitude inverted from the available liquefaction evidence is: (i) more uncertain than a 655 magnitude inverted from the available MMI evidence; and (ii) more uncertain than past studies have 656 reported. The former is attributable to there being far more MMI data than liquefaction data. The latter, as 657 previously discussed, is attributable to published uncertainty bounds (e.g., "M<sub>w</sub>6.8-M<sub>w</sub>7.0") being ranges 658 of the *median* magnitude considering one source of uncertainty, which is distinctly different from a CDF 659 660 of magnitude. Consequently, these results suggest more uncertainty than is adopted in the NSHMP. Magnitudes above M<sub>w</sub>6.7 and below M<sub>w</sub>7.5, for example (i.e., the limits of the NSHMP weighting), have 661 non-trivial probabilities of producing the field evidence. It should also be noted that these results do not 662 663 include uncertainty within the aging-correction models (which are certainly uncertain). That is, the models provide a median correction factor. To include the epistemic uncertainty of model selection, each CDF was 664 665 weighted to produce an ensemble. As with the MMI results, our selection of weights is based more on judgement than on quantitative evidence. While we agree with the developers of aging-correction models 666 that corrections are likely warranted, these corrections are actively debated and likely have large 667 668 uncertainty. The weights selected and their justifications are as follows. Model C (0.5) is the only site-669 specific method. It uses detailed subsurface geotechnical measurements from each study site to compute 670 site-specific corrections and thus arguably warrants the greatest weighting. In contrast, Models A (0.25)671 and B (0.25) both rely on a sampling of the median characteristics of the geologic unit in which each study site resides (e.g., the age of the unit). These models, which do not use site-specific information, are 672 673 therefore better suited for regional scale analyses (i.e., where subsurface data is unavailable). Nonetheless, 674 we argue these models warrant weighting given the overall uncertainty and debate surrounding aging 675 corrections. The ensemble CDF has a median of  $\sim M_w 6.90$  and 95% CI of  $M_w 6.05$  to  $M_w 7.85$ .



**Figure 13.** CDFs of rupture magnitude, as inverted from liquefaction observations using three age correction models, wherein seismic sources are modelled as points. NSHMP = National Seismic Hazard Model Project (Petersen et al., 2014, 2020).

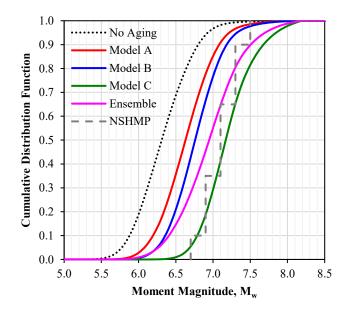
## 676 *5.5 Analysis of Liquefaction Data: Fault Search (Seismic Sources Modelled as Faults)*

To determine whether the liquefaction data can constrain the seismic source beyond a point location, the 677 678 epicenters were next converted to faults. The discretization and treatment of hypothetical fault sources was 679 identical to the MMI analyses. The results without correction for soil aging are shown in Figure 14. It can 680 be seen that the data are incapable of constraining the seismic source to the Woodstock Fault, or to any 681 other hypothetical alignment. The faults deemed most likely to produce the field evidence strike SE (like 682 the predominant result from MMI analysis), but faults with nearly any orientation also have high relative 683 likelihoods. These faults all strike through the general area previously identified via the epicenter search 684 (i.e., through the northern Charleston Peninsula). While faults striking NE do have some likelihood of 685 producing the evidence, these faults deviate from the proposed position of the Woodstock Fault. Results using each of the aging correction models are very similar to those in Figure 14 and are therefore not 686 687 presented. The use of these models increases the inverted magnitudes (and thus increases the fault lengths in Figure 14), but otherwise has little effect on the inferred fault location or orientation. As with the MMI 688 689 analyses, this should not be interpreted to mean that the Woodstock Fault or any similar alignment was not 690 the source of the 1886 earthquake. Rather, this should be interpreted to mean that many aspects of the causative fault cannot be constrained with the available macroseismic evidence. This evidence is only 691 marginally supportive of the Woodstock Fault's existence, which is not to say it doesn't exist. This 692 693 conclusion might be different if more liquefaction study sites were available to the analysis, or if the 694 component prediction models were less uncertain.



**Figure 14.** Likelihoods of hypothetical faults with differing orientations and lengths to produce the liquefaction evidence (no aging correction). Black rectangles = Woodstock fault projection (Durá-Gómez and Talwani 2009a,b); tan circles = historical epicenters ( $M_w$ >1.0; 1900-present) (USGS, 2022).

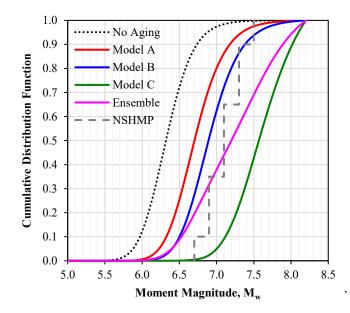
Shown in Figure 15 are the CDFs of M<sub>w</sub> inverted from liquefaction data via the "fault search." The 695 696 large uncertainty of a magnitude inferred from the liquefaction data is again apparent, as is the influence of 697 soil aging and its correction. It is again emphasized that these results do not include every source of 698 uncertainty. Omitted, for example, are the uncertainties of: (i) fault depth and length; (ii) fault dip; (iii) site  $V_{S30}$ ; and (iv) aging-correction model uncertainty. For each of these inputs only a median prediction was 699 700 used. To account for the uncertainty of selecting aging-correction models, each CDF was weighted per the prior scheme to produce an ensemble CDF, as shown in Figure 15. This ensemble has a median of M<sub>w</sub>6.95 701 702 and 95% CI of M<sub>w</sub>6.15 to M<sub>w</sub>7.85, suggesting a similar median magnitude as the NSHMP (Petersen et al., 703 2014, 2020) but much greater uncertainty.



**Figure 15.** CDFs of rupture magnitude, as inverted from liquefaction observations using three age correction models, wherein seismic sources are modelled as faults. NSHMP = National Seismic Hazard Model Project (Petersen et al., 2014, 2020).

## 704 5.6 Analysis of Liquefaction Data: Woodstock Fault

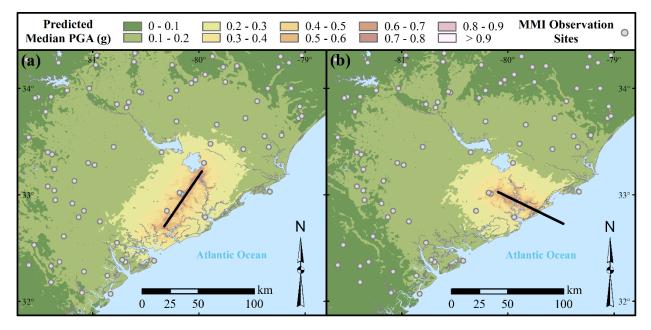
705 Finally, the preceding analyses were repeated, but conditioned solely on the Woodstock Fault as hypothesized by Durá-Gómez and Talwani (2009a,b). The computed CDFs of M<sub>w</sub>, including the weighted 706 707 ensemble, are shown in Figure 16. Our preferred CDF (i.e., the ensemble) has a median of  $M_w7.20$  and 95% 708 CI of M<sub>w</sub>6.30 to M<sub>w</sub>8.10. Relative to the CDF conditioned on an unknown source in Figure 15, conditioning 709 on the Woodstock Fault increases the inverted magnitude by  $\sim 0.2 M_w$  and increases the  $M_w$  uncertainty. 710 This counterintuitive increase in uncertainty is attributable to the epistemic uncertainty of selecting aging-711 correction models, which is observed to increase with increasing rupture magnitude. Because conditioning on the Woodstock Fault increases the inverted rupture magnitude, differences between the aging-correction 712 713 models become more apparent. Nonetheless, and similar to the analysis of MMI data, the uncertainties of 714 fault location and geometry appear relatively minor compared to other uncertainties, given the data 715 available for analysis. As shown in Figure 14, faults with diverse orientations were found to have similar 716 likelihoods of producing the liquefaction evidence. In other words, the data are sufficient to constrain the 717 source to a general area but are insufficient to constrain the source to a specific fault alignment. However, 718 because similar magnitudes are inferred for these faults irrespective of their positions, the CDFs conditioned 719 on the Woodstock Fault and the CDFs conditioned on an unknown source are similar.



**Figure 16.** CDFs of rupture magnitude, as inverted from liquefaction observations using three age correction models and assuming the Woodstock Fault (Durá-Gómez and Talwani, 2009a,b) to be the earthquake source. NSHMP = National Seismic Hazard Model Project (Petersen et al., 2014, 2020).

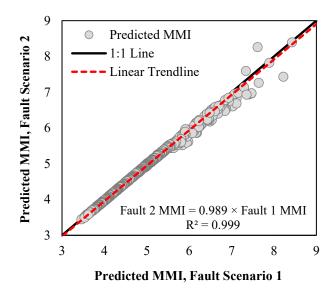
### 720 *5.7 Forward Predictions of an 1886-Like Rupture*

Based on the preceding, which indicated that the intensity and liquefaction data alone cannot well constrain 721 722 the source of the 1886 event beyond a general region, we forward predict the median PGAs expected in an 723 "1886-like" rupture. As part of this effort, two fault locations and three magnitudes are considered. The 724 purpose of these predictions is to: (i) further illustrate why constraint of the source model is difficult, given 725 data limitations; and (ii) provide predictions for this scenario event using the latest predictive models, which 726 may be useful for regional-scale consequence modelling or science communication. Shown in Figure 17, 727 for example, are predictions considering an M<sub>w</sub>7.1 rupture and two strike-slip fault scenarios: (i) fault 1, 728 which is similar to the Woodstock fault; and (ii) fault 2, which is further east, perpendicular to fault 1, and 729 similar to hypothetical faults that were shown to have high likelihoods of producing the MMI and 730 liquefaction evidence (see Figures 9 and 14). Among previously hypothesized sources of the 1886 731 earthquake, fault 2 is most similar to the Deer Park lineament proposed by Marple and Hurd (2020). Although the predicted PGAs do differ in the nearfield (e.g., within 50 km of the source), the overall 732 expectation at regional scale is obviously similar. Importantly, and as previously discussed, the Bakun et 733 734 al. (2002) dataset does not contain an abundance of 1886 MMI observations in the vicinity of Charleston, 735 or in the near field more generally. Only 16 observations, for example, are within 100 km of Charleston. 736 As a result, the predicted PGAs, and by corollary MMIs, are relatively similar at most observation sites.



**Figure 17.** Predictions of median PGA in the Charleston region, as computed using 17 NGA-East GMMs (Goulet et al., 2018) and the Harmon et al. (2019) site-response model, and considering an M<sub>w</sub>7.1 rupture on: (a) fault scenario 1; and (b) fault scenario 2, as described in the text.

Plotted in Figure 18, considering these two fault scenarios, are the predicted MMIs at each of the intensity report locations in the Bakun et al. (2002) dataset. For this example, the Atkinson and Kaka (2007) IM-MMI model was adopted. It can be seen that large differences between the predicted MMIs (say, > 0.5 MMI) are rare. In turn, the difficulty of source-model constraint is apparent, given that the two scenarios lead to perceptible differences at relatively few observation sites. While an analogous plot at sites of liquefaction evidence would show larger differences between expected PGAs (given that all such sites are in the near field), the large finite-sample uncertainty of this smaller dataset leads to a similar outcome, with a variety of faults having relatively high likelihoods of producing the evidence (see Figure 14). It follows that additional near-field evidence (whether MMI or liquefaction) could be especially influential to future studies of the 1886 earthquake. Following the same approach, predictions were made for Mw6.6 and Mw7.6 ruptures. These results are shown in the electronic supplement, where each of the forward predictions is also provided as a GIS map package.



**Figure 18.** Predictions of MMI using the Atkinson and Kaka (2007) IM-MMI model for fault scenarios 1 and 2 with rupture magnitude  $M_w7.1$ . Black line = 1:1 line; red dashed line = linear trendline.

## 737 *5.8 Limitations and Uncertainties*

738 The findings presented herein are inherently tied to the field evidence currently available and adopted for analysis. Reinterpretation or augmentation of these observations, which we assume to be independent 739 740 events, would potentially change these findings, as would the adoption of new component models (e.g., to 741 predict ground motions, site response, MMI, liquefaction triggering, or liquefaction manifestation). Undoubtedly, the modeling of these phenomena will continually advance, warranting future analyses of the 742 743 1886 Charleston earthquake. ENA ground-motion modeling has advanced, for example, yet GMMs remain especially uncertain and untested at large magnitudes and will continue to evolve. And, as a supplement to 744 745 empirical GMMs, physics-based ground motion simulations could provide new insights into the 1886 746 rupture via more explicit modeling of influential site, path, and source effects (e.g., stress drop), as has been 747 shown for other historical earthquakes (e.g., Lozos, 2016). While empirical GMMs are ubiquitous in 748 earthquake science/engineering and implicitly account for many salient effects, a more explicit accounting 749 could help to reduce uncertainty. Moreover, the goal of this study was to assess the degree to which 750 macroseismic evidence can constrain the 1886 source model, rather than to perform a broad investigation 751 of all geophysical and modern seismological evidence, which would fall under the purview of other 752 investigators with different expertise. In addition, it should be noted that many, but not all, sources of 753 uncertainty were accounted for in the analyses. Neglected, for example, were the uncertainties of site 754 characterization (i.e.,  $V_s 30$ ) and site observations (i.e., MMI and liquefaction responses). The inclusion of 755 these and other uncertainties could potentially broaden the  $M_w$  CDFs computed herein. It must similarly be emphasized that a thorough reinterpretation of the more than 1000 original intensity reports, to include the 756

assignment of site-specific uncertainties, bias corrections, and/or weighting schemes, could potentially

change the results we present. And, as previously emphasized, our ensemble M<sub>w</sub> CDFs utilize judgement-

based weighting schemes. While justifications were provided for these weights, readers might argue for

other weights, and thus draw other conclusions from our suites of  $M_w$  CDFs. Ultimately, future analyses

will confirm or revise the conclusions reached in this study and summarized below.

#### 762 6. Conclusions

763 Prior studies of MMI and liquefaction data resulting from the 1886 Charleston, SC, earthquake have several 764 limitations. Namely, these studies tend to: (i) either be deterministic or account for uncertainties in an informal manner (e.g., it is often unclear what published uncertainty bounds represent and which 765 766 uncertainties are, and are not, included); (ii) assume that the 1886 event was caused by a particular fault 767 (i.e., the Woodstock Fault) without investigating the uncertainty of this assumption or the ability of the field 768 data to constrain source traits beyond magnitude (i.e., fault location, geometry); and (iii) rely on models for 769 predicting various phenomena (e.g., ground motions, site response, liquefaction response, MMI) that have 770 since been superseded or augmented (e.g., by the NGA East project's 17 GMMs). Accordingly, this study 771 presented probabilistic seismic-source inversions of the 1886 earthquake using the latest predictive models 772 and a novel inversion methodology. With this approach, the likelihood of a rupture with some magnitude, 773 location, and geometry to produce a set of field observations is computed. Repeating for enumerable 774 hypothetical sources results in a regional scale constraint of the likely source traits, to the extent that 775 observational data permits. With this approach, magnitude probability distributions conditioned on both an 776 unknown source and on the hypothesized Woodstock Fault (Durá-Gómez and Talwani, 2009a,b) were 777 computed and compared to that used in the U.S. NSHMP (Petersen et al., 2014, 2020). The most salient 778 findings of this study, subject to the limitations and uncertainties summarized in section 5.8, are:

Neither the location nor orientation of the 1886 fault rupture could be confidently constrained by the macroseismic evidence and models utilized herein, given the large uncertainties inherent to each. Hypothetical faults in a range of locations and with various alignments were deemed to have high relative likelihoods of producing this evidence. Considering both types of evidence and all analyses, faults striking SE and partly offshore were predominantly identified as having the greatest likelihood. Yet these faults ranged in location and other faults, with very different strikes, were also often found to be relatively likely.

Collectively, the evidence does not provide strong support for the hypothesized Woodstock Fault.
 One analysis (that of MMI data using the WEA12 IM-MMI model) found a NE-striking "Woodstock
 like" fault to be the most likely source of the field evidence. Yet most analyses – in so far as
 supporting the Woodstock Fault – point to the likelihood of a seismic source somewhere north of

Charleston but deem the Woodstock Fault to be relatively unlikely. This is not to say the fault does not exist (geophysical investigations and modern seismological data clearly suggest that many faults exist in the area), but rather, that many aspects of the 1886 source model cannot be well constrained with the available macroseismic evidence and models. This result might change if: (i) more MMI and liquefaction evidence were available – particularly in the near field; (ii) the MMI evidence was reinterpreted to remove or correct outliers, and to assign observation-specific uncertainties; or (iii) if the various required component models were less uncertain.

- In the absence of these developments, certain aspects of the 1886 fault rupture can only be constrained
   with other seismologic, geologic, and/or coseismic data and interpretation.
- When conditioned on the Woodstock Fault proposed by Durá-Gómez and Talwani (2009a,b), and generally supported by others, our preferred CDF of M<sub>w</sub> inverted from MMI data has a median of M<sub>w</sub>7.10 and 95% CI of M<sub>w</sub>7.0 to M<sub>w</sub>7.75. Of all uncertainties considered, the epistemic uncertainty of IM-MMI model selection was larger than any other, since different models may give significantly different MMI predictions for a given IM. In this regard, efforts to better quantify the suitability of IM-MMI models to the study region, and in turn, to refine the weighting scheme used herein, could have a significant influence on our overall conclusions.
- The results from MMI analysis show some site-to-source distance bias, with magnitudes inverted
   from more distal MMI observations tending to be larger. Possible reasons for this bias were discussed
   but could be further investigated in the future.
- When conditioned on the Woodstock Fault, our preferred CDF of M<sub>w</sub> inverted from liquefaction data has a median of M<sub>w</sub>7.20 and 95% CI of M<sub>w</sub>6.30 to M<sub>w</sub>8.10. The greater M<sub>w</sub> uncertainty, relative to that from MMI analysis, is attributable to there being fewer liquefaction study sites, the compilation of which requires both an observation of liquefaction response *and* subsurface geotechnical testing. In addition, the uncertainties of whether and how to correct for soil-aging effects considerably augmented the uncertainty. As such, efforts to quantify the suitability and uncertainties of aging-correction models could have significant influence on our overall conclusions.
- When conditioned on an unknown seismic source, the CDFs of  $M_{\rm w}$  inverted from MMI and 816 • liquefaction data did not differ greatly from the CDFs conditioned on the Woodstock Fault. Our 817 interpretation is that while faults with a range of locations and alignments were found similarly likely 818 to produce the field evidence, these faults were inferred to have mostly similar magnitude 819 820 distributions. Thus, the results of this study pertaining to the magnitude of the 1886 rupture would 821 not necessarily change if the rupture's exact position was known. This should not be interpreted to 822 mean that the uncertainties of source location and geometry are inconsequential to inverse-analyses 823 of macroseismic data. As demonstrated via simulated inversions of modern earthquakes (Rasanen

and Maurer, 2021, 2022), these uncertainties are often considerable. This is especially the case, for example, when the effects of a distant, large  $M_w$  rupture cannot be distinguished from the effects of a nearby, small  $M_w$  rupture. In such cases, constraint of the source location can dramatically reduce the overall uncertainty of the inverted magnitude.

- 828 Collectively, the results largely support the M<sub>w</sub> distribution adopted by the NSHMP, which ranges from  $M_w 6.7$  to  $M_w 7.5$  with a median of  $M_w 7.1$ . While analyses indicate that  $M_w < 6.7$  ruptures have 829 830 potential to produce the observed liquefaction response (particularly when the uncertainty of selecting 831 an aging-correction model is considered), the MMI evidence suggests a near-zero likelihood of such magnitudes. Conversely, both types of evidence suggest that  $M_w > 7.5$  events have potential to 832 produce the field evidence. While this conclusion hinges on which component models are adopted to 833 predict soil-aging effects and MMI, we fail to find conclusive evidence for outright rejecting some 834 835 models in favor of others. In the absence of such evidence, the possibility of  $M_w > 7.5$  ruptures would 836 merit consideration.
- Ultimately, the flexible inversion methodology employed herein is not specific to ENA, or to certain
   types of macroseismic evidence, but rather is applicable to any seismic zone and to any co-seismic
   response for which probabilistic prediction models exist. This methodology allows for uncertainty to
   be accounted for in a more complete and transparent manner when inverting seismic parameters from
   macroseismic evidence.

#### 842 Declaration of competing interests

843 The authors declare no competing interests.

## 844 Acknowledgements

The presented study is based on work supported by the National Science Foundation (NSF) under Grant No. CMMI-1751216 and by the NSF Graduate Research Fellowship Program under Grant No. DGE-1762114. However, any opinions, findings, and conclusions or recommendations expressed in this paper are those of the authors and do not necessarily reflect the views of NSF. We also gratefully acknowledge Silvia Mazzoni of the John Garrick Institute for Risk Sciences at UCLA for developing NGA East ground motion characterization tools based on the results of Goulet et al. (2018), which were utilized in this study.

#### 851 Appendix A. Supplementary data

All data and models utilized in this study are publicly available, as cited herein. The supplemental material includes 8 additional figures that pertain to the analysis of MMI data using the "epicenter search" and "fault search" methods. The results presented in these figures are summarized and discussed in 5.1-5.2. Also

- 855 included are maps (figures and GIS files) of the predicted median PGAs in an "1886-like" event,
- 856 considering three magnitudes and two fault alignments.

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