

Fast Probabilistic Seismic Hazard Analysis through Adaptive Importance Sampling

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ABSTRACT

Probabilistic Seismic Hazard Analysis (PSHA) traditionally relies on two computationally intensive approaches: (a) Riemann Sum and (b) conventional Monte Carlo (MC) integration. The former requires fine slices across magnitude, distance, and ground motion, and the latter demands extensive synthetic earthquake catalogs. Both approaches become notably resource-intensive for low-probability seismic hazards, where achieving a COV of 1% for a 10^{-4} annual hazard probability may require 10^8 MC samples. We introduce Adaptive Importance Sampling (AIS) PSHA, a novel framework to approximate optimal importance sampling (IS) distributions and dramatically reduce the number of MC samples to estimate hazards. We evaluate the efficiency and accuracy of our proposed framework using Pacific Earthquake Engineering Research Center (PEER) PSHA benchmarks that cover various seismic sources, including areal, vertical, and dipping faults, as well as combined types. Our approach computes seismic hazard up to 3.7×10^4 and 7.1×10^3 times faster than Riemann Sum and traditional MC methods, respectively, maintaining COVs below 1%. We also propose an enhanced approach with a “smart” AIS PSHA variant that leverages the sampling densities from similar ground motion intensities. This variant outperforms even “smart” implementations of Riemann Sum with enhanced grid discretizations by a factor of up to 130. Moreover, we demonstrate theoretically that optimal IS distributions are equivalent to hazard disaggregation distributions. Empirically, we show the approximated optimal IS and the disaggregation distributions are closely alike, e.g., with a Kolmogorov–Smirnov statistic between 0.017 and 0.113. This approach is broadly applicable, especially for PSHA cases requiring extensive logic trees and epistemic uncertainty.

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KEY POINTS

- The Adaptive Importance Sampling (AIS) is introduced for fast PSHA computations.
- AIS PSHA is significantly faster than traditional numerical approaches and easily facilitates disaggregation.
- The approach is broadly applicable to PSHA involving extensive logic trees.

Supplemental Material

INTRODUCTION

Probabilistic Seismic Hazard Analysis (PSHA) has become a foundational method for determining seismic design levels and conducting regional seismic risk analyses since its first inception (Cornell, 1968; U. S. Nuclear Regulatory Commission, 2007; McGuire, 2008; ASCE, 2022; Kennedy et al., 1980; Ceferino et al., 2020; Silva et al., 2020; Baker et al., 2021; Papadopoulos and Bazzurro, 2021; Arora and Ceferino, 2023). Since we cannot solve PSHA analytically due to the complexity in seismic source and ground motion models, numerous researchers have developed computer software for PSHA computation (Cornell, 1968; McGuire, 1976; Kiremidjian et al., 1982; Field et al., 2003; Ordaz et al., 2013; Pagani et al., 2014). The existing software primarily employs Riemann Sum for numerical integration of PSHA (Thomas et al., 2010; Hale et al., 2018). The Riemann Sum offers robust PSHA integration with sufficiently dense grids. However, this method generally incurs a significant computational load exponentially increasing with the number of grids and dimensions in multi-dimensional integrations (Philippe and Robert, 2001). Furthermore, the results are highly sensitive to the chosen grid design, especially for low exceedance probabilities (Thomas et al., 2010; Hale et al., 2018). Alternatively, other software adopted Monte-Carlo (MC) integration for PSHA (Assatourians and Atkinson, 2013, 2019). MC integration calculates exceedance probabilities by generating random synthetic earthquake catalogs (Musson, 2000). MC's primary advantage lies in its straightforward concept compared to Riemann Sum (Musson, 2000; Dick et al., 2013). Nevertheless, MC framework requires a substantially-long synthetic catalog to accurately estimate hazards from rare events, especially for low exceedance probabilities, e.g., $p < 10^{-4}$ /yr (Kroese et al., 2014).

Importance Sampling (IS) can offer a solution to this rare event simulation (Tokdar and Kass, 2010). IS was initially introduced in statistical physics (Hammersley and Morton, 1954) to improve the computational efficiency of rare event simulation. IS relies on identifying an appropriate probability distribution ("IS distribution") to explore low-probability spaces effectively. Researchers use the IS distribution to sample rare events with a higher likelihood than conventional MC and then correct their frequency through weights, significantly reducing the number of samples to compute low probabilities (Robert et al., 1999). However, finding such an appropriate distribution can be challenging because the distribution differs for each problem being solved. Thus, many numerical experiments are often conducted first through trial-and-error to identify IS distributions, which is still computationally expensive.

46 In regional seismic risk analysis, numerous studies have been conducted to sample hazard-consistent earthquake ground
47 motions (Crowley and Bommer, 2006; Kiremidjian et al., 2007; Jayaram and Baker, 2010; Han and Davidson, 2012; Manzour
48 et al., 2016; Christou et al., 2018; Kavvada et al., 2022). Kiremidjian et al. (2007) first introduced IS distributions that sample
49 large-magnitude earthquakes with a high probability to reduce the computational burden of seismic hazard and risk analyses.
50 Jayaram and Baker (2010) expanded the approach by defining IS distributions to sample high-intensity ground motions.
51 However, Jayaram and Baker (2010) highlighted the computational challenges to identify an effective IS distribution and
52 ended up using K-mean clustering to reduce the number of ground motion samples. Rahimi and Mahsuli (2019) applied
53 the IS sampling density as the normal distribution centered at the “design point” derived from the first- and second-order
54 reliability method to PSHA computation.

55 To find effective IS distributions, computational statisticians have developed a general framework to find them through
56 iterative algorithms denominated “Adaptive Importance Sampling (AIS)” (Bugallo et al., 2017). Various AIS algorithms
57 have been introduced, e.g., cross-entropy based AIS, Vegas, Divonne, and Miser (Lepage, 1978; Friedman and Wright, 1981;
58 Rubinstein, 1997; Press and Farrar, 1990; Rubinstein and Kroese, 2004; Bugallo et al., 2017). The application of AIS is widely
59 adopted to solve the integration with high dimensions, such as in the field of statistical physics, finance, reliability engi-
60 neering, and signal processing (Au and Beck, 2001; Kappen and Ruiz, 2016; Nieto and Ruiz, 2016; Bugallo et al., 2017).
61 Although previous studies suggest the use of IS for PSHA calculation (Jayaram and Baker, 2010; Rahimi and Mahsuli, 2019),
62 no research has been published regarding the use of AIS for this purpose. AIS can provide a general methodology for identi-
63 fying an appropriate IS distribution for seismic hazards, eliminating computationally expensive experiments in the regular
64 IS approach, whose efficiency typically depends on the researcher’s experience in the field.

65 Among various AIS algorithms (Lepage, 1978; Friedman and Wright, 1981; Rubinstein, 1997; Press and Farrar, 1990;
66 Rubinstein and Kroese, 2004; Bugallo et al., 2017), in this study, we apply the VEGAS algorithm for PSHA to leverage its
67 straightforward mathematical framework and fast convergence. Thus, this study introduces a novel computational method
68 for PSHA curve calculation using the AIS VEGAS algorithm (Lepage, 1978, 2021). This paper also shows that AIS facilitates
69 hazard disaggregation, the relative contribution of each random variable to the overall hazard (Bazzurro and Cornell, 1999).

70 We explore three key aspects of the method: 1) the enhancement of computational efficiency and accuracy that the AIS
71 algorithm offers over traditional methods; 2) the process of obtaining hazard disaggregation through AIS; and 3) the presen-
72 tation of “smart” AIS to further improve the performance of the algorithm. We present the theoretical background of AIS
73 PSHA and validate the method through numerical examples.

74 **MATHEMATICAL FORMULATION**

75 **Probabilistic Seismic Hazard Analysis (PSHA)**

76 At a site of interest, the annual rate of ground motion exceedance from a single source can be calculated as:

$$\lambda(X > a) = \nu \int_{\varepsilon_{\min}}^{\varepsilon_{\max}} \int_{r_{\min}}^{r_{\max}} \int_{m_{\min}}^{m_{\max}} I(X > a|m, r, \varepsilon) f_{M,R,\varepsilon}(m, r, \varepsilon) dm dr d\varepsilon \quad (1)$$

77 , where $\lambda(X > a)$ is the annual rate that ground motion, X , exceeds the target ground motion intensity, a , e.g., peak ground
 78 acceleration. ν is the annual rate of earthquake occurrence with magnitude greater than m_{\min} from the source, M is the
 79 earthquake magnitude, m_{\min} and m_{\max} are minimum and maximum magnitudes considered for the source, R is the source-
 80 to-site distance, r_{\min} and r_{\max} are minimum and maximum source-site distances, ε is a standard normal random variable
 81 for generating earthquake ground motion, ε_{\min} and ε_{\max} are ε 's minimum and maximum values for PSHA computations
 82 (generally, $\varepsilon_{\max} \geq 6$ and $\varepsilon_{\min} \leq -6$; [Bommer and Abrahamson \(2006\)](#)), $f_{M,R,\varepsilon}(m, r, \varepsilon)$ is joint probability density function
 83 (PDF) of M , R , and ε , $I(X > a|m, r, \varepsilon)$ is the indicator function that takes 1 when $X > a$, otherwise, 0. The ground motion X
 84 given M , R , and ε is generally calculated using ground motion models ([Bozorgnia et al., 2014](#); [Goulet et al., 2021](#)). The models
 85 usually assume the log-normal distribution for the ground motion given explanatory variables such as M and R . Naturally,
 86 these models provide the mean and standard deviation of logarithmic ground motion. Thus, the random ground motion is
 87 calculated as:

$$\log X = \mu(M, R) + \varepsilon\sigma(M, R) \quad (2)$$

88 , where μ and σ are the mean and standard deviation of logarithmic earthquake ground motion.

89 If we assume that the ground motion random variable ε is independent with respect to the M and R ([McGuire, 1995](#)),
 90 Eq. (1) can be modified as

$$\lambda(X > a) = \nu \int_{\varepsilon_{\min}}^{\varepsilon_{\max}} \int_{r_{\min}}^{r_{\max}} \int_{m_{\min}}^{m_{\max}} I(x > a|m, r, \varepsilon) f_{\varepsilon}(\varepsilon) f_{R|M}(r|m) f_M(m) dm dr d\varepsilon \quad (3)$$

91 , where $f_M(m)$, $f_{R|M}(r|m)$, $f_{\varepsilon}(\varepsilon)$ are PDF of M , R given M , and ε . Under the point source assumption, the distance R and
 92 magnitude M become independent random variables. Thus, the seismic hazard is given by

$$\lambda(X > a) = \nu \int_{\varepsilon_{\min}}^{\varepsilon_{\max}} \int_{r_{\min}}^{r_{\max}} \int_{m_{\min}}^{m_{\max}} I(x > a|m, r, \varepsilon) f_{\varepsilon}(\varepsilon) f_R(r) f_M(m) dm dr d\varepsilon \quad (4)$$

93 The total seismic hazard from multiple seismic sources (e.g., different faults) is the sum of each, $\Lambda(X > a) = \sum_{i=1}^{n_s} \lambda_i(X > a)$,
 94 where $\Lambda(X > a)$ is the total annual rate of exceedance of ground motion, a , i is index for seismic sources, and n_s is the
 95 total number of seismic sources. Under the assumption of the Poisson process, the annual probability of exceedance can be
 96 converted from the annual rate of exceedance by $1 - e^{-\Lambda(X > a)}$.

97 Hazard disaggregation

98 We can also disaggregate the hazard to better understand the earthquakes that contribute most to the hazard. Disaggregation
 99 is also used to select the input seismic records and conduct non-linear time-history analyses for the design of many critical
 100 buildings (Bazzurro and Cornell, 1999; U. S. Nuclear Regulatory Commission, 2007). Mathematically, disaggregation of the
 101 hazard is the joint probability distribution of the M , R , and \mathcal{E} conditional on different levels of hazards a to quantify the
 102 contributions of each component. The disaggregation of PSHA can be formulated using Bayes' theorem as

$$\begin{aligned}
 P(M, R, \mathcal{E}|X > a) &= \frac{P(X > a \cap M, R, \mathcal{E})}{P(X > a)} \\
 &= \frac{I(X > a|M, R, \mathcal{E})P(M, R, \mathcal{E})}{\sum_M \sum_R \sum_{\mathcal{E}} I(X > a|M, R, \mathcal{E})P(M, R, \mathcal{E})}
 \end{aligned} \tag{5}$$

103 , where $P(X > a|M, R, \mathcal{E})$ is the probability of ground motion X is greater than a given M , R , and \mathcal{E} , $P(M, R, \mathcal{E})$ is joint probabil-
 104 ity of M , R , and \mathcal{E} , and $P(X > a)$ is the total probability that the ground motion is greater than a , which is $\sum_M \sum_R \sum_{\mathcal{E}} P(X >$
 105 $a|M, R, \mathcal{E})P(M, R, \mathcal{E})$. Note that $P(X > a|M, R, \mathcal{E})$ can be expressed as an indicator function, $I(X > a|m, r, \varepsilon)$, because the
 106 probability of ground motion X greater than a can only be 1 or 0 given m , r , and ε (see Eq. (2)).

107 By replacing probability mass function, $P(M, R, \mathcal{E})$, with probability density function, $f_{M,R,\mathcal{E}}(m, r, \varepsilon)$, and changing the
 108 summation into integration, Eq. (5) can be expressed as:

$$f(m, r, \varepsilon|X > a) = \frac{I(X > a|m, r, \varepsilon)f_{M,R,\mathcal{E}}(m, r, \varepsilon)}{\iiint I(X > a|m, r, \varepsilon)f_{M,R,\mathcal{E}}(m, r, \varepsilon)dmr d\varepsilon} \tag{6}$$

109 By Eq. (1), the denominator of Eq. (6) equals λ/ν . Thus,

$$f(m, r, \varepsilon|X > a) = \nu \cdot \frac{I(X > a|m, r, \varepsilon)f_{M,R,\mathcal{E}}(m, r, \varepsilon)}{\lambda} \tag{7}$$

110 CURRENT NUMERICAL SOLUTIONS

111 **The Riemann Sum** This method computes PSHA curves by summing the areas of partitioned (m, r, ε) cuboids. The
 112 Riemann Sum for Eq. (1) can be expressed as:

$$\lambda(X > a) = \nu \sum_{k=1}^{N_{\varepsilon}} \sum_{j=1}^{N_r} \sum_{i=1}^{N_m} I(x > a|m_i, r_j, \varepsilon_k) f_{M,R,\mathcal{E}}(m_i, r_j, \varepsilon_k) \Delta m \Delta r \Delta \varepsilon \tag{8}$$

113 , where Δm , Δr , and $\Delta \varepsilon$ are grid step size for Riemann Sum, N_m , N_r , and N_{ε} are the total number of grids, satisfying $N_x \Delta x$
 114 $= x_{\max} - x_{\min}$, where $x = \{m, r, \varepsilon\}$. Note that $f_{M,R,\mathcal{E}}(m_i, r_j, \varepsilon_k) \Delta m \Delta r \Delta \varepsilon$ is equivalent to the probability at (m, r, ε) such that
 115 $|m - m_i| < \Delta m/2$, $|r - r_j| < \Delta r/2$, and $|\varepsilon - \varepsilon_k| < \Delta \varepsilon/2$. The accuracy of the Riemann Sum depends on the grid size. Utilizing
 116 finer grids enhances the accuracy of the summation. However, the computation time is inversely proportional to the grid step

TABLE 1
Comparison of time complexity of various PSHA algorithms

Algorithm	N_m	N_r	N_ε	N_s	N_a
Riemann Sum	N	N	$N (1^*)$	-	1
Conventional MC	1	1	1	N	1
IS MC	1	1	1	N	1
VEGAS AIS (this study)	1	1	1	N	$N (< N^\dagger)$

N_m , the number of magnitude grids; N_r , the number of distance grids; N_ε , the number of ε grids; N_s , the number of (m, r, ε) samples; N_a , the number of ground motion intensity of interest (e.g., 0.1 g)

* when Eq. (26) is used

† when “smart” AIS is used

117 size and thus proportional to the number of grids. For three-dimensional PSHA summation, the computation time scales with
 118 $N_m \times N_r \times N_\varepsilon$ (Table 1). Notably, since the distance PDF ($f_R(r)$) cannot be analytically determined in practice, integration
 119 often extends over latitude (ϕ), longitude (ψ), and depth (z), increasing the dimensions from three (m, r, ε) to five ($m, \phi, \psi,$
 120 z, ε). Therefore, the Riemann Sum for seismic hazard becomes even more computationally intensive, a phenomenon known
 121 as “the curse of dimensionality” due to the exponential increase in computation time with the number of dimensions (Novak
 122 and Ritter, 1997).

123 **Conventional Monte-Carlo (MC)** This method simulates many synthetic earthquake ground motions and calculates PSHA
 124 by assessing the frequency with which ground motion intensities exceed a certain threshold. MC PSHA is computed as

$$\begin{aligned}\hat{\lambda}(X > a) &= \frac{\nu}{N_s} \sum_{i=1}^{N_s} I(X_i > a | M_i, R_i, \mathcal{E}_i) \\ &= \frac{1}{T} \sum_{i=1}^{N_s} I(X_i > a | M_i, R_i, \mathcal{E}_i)\end{aligned}\quad (9)$$

125 , where X_i denotes the simulated ground motions, and N_s is the total number of samples. M_i, R_i, \mathcal{E}_i are random samples from
 126 $f_{M,R,\mathcal{E}}(m, r, \varepsilon)$, T represents the equivalent catalog duration equal to N_s/ν . MC PSHA is unbiased, i.e., the expectation of
 127 $\hat{\lambda}(X > a)$ is the same as λ , and the variance of $\hat{\lambda}$, $\text{VAR}[\hat{\lambda}]$, is $(\nu\lambda - \lambda^2)/N_s$ (see Appendix A for derivation). Note that $\text{VAR}[\hat{\lambda}]$
 128 is always positive since $\nu \geq \lambda$. The coefficient of variation, COV, a widely used metric to estimate relative uncertainty with
 129 respect to the mean, is:

$$\text{COV} = \frac{\sqrt{\text{VAR}[\hat{\lambda}]}}{E[\hat{\lambda}]} = \sqrt{\frac{\nu - \lambda}{N_s \lambda}} \quad (10)$$

130 COV is inversely proportional to the square root of the number of MC samples (N_s), indicating that increasing N_s naturally
 131 improves the estimate's accuracy. The target exceedance rate, λ , also affects the MC accuracy. The lower λ leads to poor
 132 accuracy with fixed N_s and ν . Thus, a sufficiently long earthquake catalog is required to accurately estimate the event from a
 133 long return period. It is also notable that COV is increasing with earthquake occurrence rate, ν , indicating that MC PSHA is
 134 a more challenging task in regions with higher earthquake activity. This can also be explained intuitively. Regions with high
 135 seismic activity experience more earthquakes per year than less active areas; therefore, when calculating the ground motion
 136 exceedance probability over the same period, more earthquakes need to be considered in these active regions.

137 In MC PSHA, each realization of the ground motion can be leveraged to compute hazard estimates at different intensity
 138 levels. An advantage of MC PSHA is that the computational time is dependent on the number of samples N_s , circumventing
 139 the inherent dimensionality problem in Riemann Sum (Table 1). In other words, the computation time of MC simulation
 140 is independent of the number of grids (N_m, N_r , and N_ε in Eq. (8)). Therefore, we could utilize fine joint probability mass
 141 functions for more precise hazard estimation without an increase in computational burden. If closed-form probability density
 142 functions are available, PSHA can be implemented without approximated discretization.

143 Conventional MC faces extreme computational challenges for low probabilities because the number of samples required
 144 to achieve low COVs dramatically increases. We can compute the required number of samples by rearranging Eq. (10):

$$N_s = \frac{1}{(\text{COV})^2} \times \frac{\nu - \lambda}{\lambda} \quad (11)$$

145 For small λ , i.e., $\nu \gg \lambda$, the Eq. (11) can be approximated as $N_s \sim \frac{1}{(COV)^2} \times \frac{\nu}{\lambda}$. To demonstrate conventional MC's extreme
 146 computational demands, we can calculate N_s for a target annual exceedance rate, λ , of 10^{-4} per year and an annual earth-
 147 quake rate $\nu = 1/\text{yr}$, which is typical values in PSHA practice (Coppersmith et al., 2014). From Eq. (11), we need $N \sim 10^8$ to
 148 achieve $COV = 1\%$.

149 Importance Sampling (IS) Integration

150 IS is a generalization of MC theory. Consider a random variable X that follows a probability density function, $f_X(x)$. The
 151 expected value of a function $u(x)$, denoted as S , is defined by

$$S = \int u(x)f_X(x)dx \quad (12)$$

152 The MC estimate is

$$\hat{S} = \frac{1}{N_s} \sum_{i=1}^{N_s} u(X_i) \quad (13)$$

153 , where X_i is sampled from $f_X(x)$.

154 By introducing an arbitrary probability density function, $q_X(x)$, Eq. (12) can be equivalently expressed as

$$S = \int u(x) \frac{f_X(x)}{q_X(x)} q_X(x) dx \quad (14)$$

155 We restrict the integration range in Eq. (14) where $f_X(x) \neq 0$ because x such that $f_X(x) = 0$ does not contribute to the inte-
 156 gration. Then, $q_X(x)$ can be any distribution with nonzero density in the integration range. Eq. (14) provides important
 157 implications in MC estimation. We can get the solution to Eq. (12) by estimating the expected value of $u(x)f_X(x)/q_X(x)$ where
 158 x follows the distribution $q_X(x)$. This approach is highly useful in numerical integration, especially when sampling from
 159 $f_X(x)$ is challenging or the population of $f_X(x)$ is extremely low in the region of importance. For example, the exceedance
 160 rate of PGA greater than 1 g is mostly contributed by the ground motion samples from large magnitude (M) earthquakes
 161 occurred at close distance (R) with large \mathcal{E} , which all typically correspond to the low probability region.

162 The IS MC estimate is:

$$\hat{S} = \frac{1}{N_s} \sum_{i=1}^{N_s} u(X_i) \frac{f_X(X_i)}{q_X(X_i)} \quad (15)$$

163 , where X_i s are sampled from $q_X(x)$, the proposed (or new) IS sampling density function. If $q_X(x)$ is equal to the original
 164 distribution, $f_X(x)$, Eq. (15) simplifies to the conventional Monte-Carlo (Eq. (13)). The ratio $f_X(x)/q_X(x)$, known as the
 165 importance weight (w_i), adjusts for the change in sampling distribution.

166 **IS PSHA** Researchers have applied IS to PSHA using different sampling functions (Jayaram and Baker, 2010; Rahimi and
 167 Mahsuli, 2019). For IS PSHA, Eq. (1) can be reformulated by introducing a new sampling joint density function, $q_{M,R,\varepsilon}(m, r, \varepsilon)$
 168 as follows

$$\lambda(X > a) = \nu \int_{\varepsilon_{\min}}^{\varepsilon_{\max}} \int_{r_{\min}}^{r_{\max}} \int_{m_{\min}}^{m_{\max}} I(x > a | m, r, \varepsilon) \frac{f_{M,R,\varepsilon}(m, r, \varepsilon)}{q_{M,R,\varepsilon}(m, r, \varepsilon)} q_{M,R,\varepsilon}(m, r, \varepsilon) dm dr d\varepsilon \quad (16)$$

169 The IS MC estimator of equation (16) is

$$\hat{\lambda}(X > a) = \frac{\nu}{N_s} \sum_{i=1}^{N_s} I(X_i > a | M_i, R_i, E_i) \frac{f_{M,R,\varepsilon}(M_i, R_i, E_i)}{q_{M,R,\varepsilon}(M_i, R_i, E_i)}$$

170 The mean of IS estimator $\hat{\lambda}$ is also unbiased like conventional MC, i.e., $E_q[\hat{\lambda}] = \lambda$ (see Appendix B for derivation). Also, the
 171 variance of the IS PSHA estimates with respect to the true λ is given by

$$\text{VAR}[\hat{\lambda}] = \frac{1}{N_s} \left(\nu^2 E_q \left[\left(I(X_i > a | M_i, R_i, E_i) \frac{f_{M,R,\varepsilon}(M_i, R_i, E_i)}{q_{M,R,\varepsilon}(M_i, R_i, E_i)} \right)^2 \right] - \lambda^2 \right) \quad (17)$$

172 (see Appendix B for derivation). Hence, the COV of IS PSHA estimate can be expressed as

$$\text{COV} = \sqrt{\frac{\nu^2 E_q \left[\left(I(X_i > a | M_i, R_i, E_i) \frac{f_{M,R,\varepsilon}(M_i, R_i, E_i)}{q_{M,R,\varepsilon}(M_i, R_i, E_i)} \right)^2 \right] - \lambda^2}{N_s \lambda^2}} \quad (18)$$

173 **Equivalency of optimal IS density and hazard disaggregation**

174 From Eq. (17), we can specifically choose a new sampling density $q_{M,R,\varepsilon}^*$ that makes $\text{VAR}[\hat{\lambda}] = 0$,

$$q_{M,R,\varepsilon}^* = \nu \frac{I(X_i > a | M_i, R_i, E_i) f_{M,R,\varepsilon}(M_i, R_i, E_i)}{\lambda} \quad (19)$$

175 Using $q_{M,R,\varepsilon}^*$, we could compute the true hazard, λ , with only one MC sample because IS MC is unbiased and the variance
 176 is zero. We call q^* the optimal IS density for PSHA calculation.

177 We note that Eq. (19) is exactly the same as Eq. (7), indicating that the optimal density, q^* , is identical to the hazard
 178 disaggregation. That is, if we find q^* , we are able to not only dramatically enhance the computational efficiency of seismic
 179 hazard estimation but also obtain the hazard disaggregation distributions as a by-product.

180 In fact, we can also see this profound relationship between the hazard and disaggregation estimates by rearranging Eq. (7):

$$\lambda \equiv \nu \cdot \frac{I(X > a | m, r, \varepsilon) f_{M,R,\varepsilon}(m, r, \varepsilon)}{f(m, r, \varepsilon | X > a)}$$

181 Because this identity holds for any values of (m, r, ε) , we can obtain the hazard at ground motion level, a , with any (m, r, ε)
182 triplet if we know $f(m, r, \varepsilon|X > a)$. The term $f_{M,R,\varepsilon}(m, r, \varepsilon)/f(m, r, \varepsilon|X > a)$ can be interpreted as the importance weight of
183 IS, and $f(m, r, \varepsilon|X > a)$, the hazard disaggregation, is the optimal density. Note that this identity also holds in the “equal”
184 hazard case, where the indicator function $I(X > a)$ becomes $I(X = a)$ (or $I(a_1 < X \leq a_2)$).

185 The disaggregation can be computationally expensive in memory usage. In Riemann Sum, we must save each sum element
186 in memory and allocate those elements into appropriate disaggregation bins, though the elements are readily computed in
187 the hazard computation process. Also, in conventional MC, we should save the long synthetic ground motion catalog with the
188 corresponding (m, r, ε) triplet to allocate those into the proper bins. These operations necessitate additional computational
189 memory and time. Therefore, finding q^* could bring significant benefits by simultaneously producing disaggregation and
190 hazard curves. However, finding q^* is not trivial in current implementations of PSHA because $I(X_i > a|M_i, R_i, E_i)$ and λ in
191 Eq. (19) are unknowns. Thus, we propose a new PSHA computation method to find it.

192 ADAPTIVE IMPORTANCE SAMPLING PSHA

193 In “adaptive” importance sampling (AIS), we iteratively train the IS density to find the optimal one (q^*) by exploring impor-
194 tant regions to compute $I(X_i > a|M_i, R_i, E_i)$ and λ with a reduced number of MC samples. Regardless of the algorithm used,
195 AIS must balance different factors in determining how many samples (N_s) should be used to train optimal IS density. If we
196 use a small number of samples to update the sampling distribution, we will not explore the important regions effectively,
197 making the number of iterations increase. On the other hand, many samples can reduce the number of iterations, however,
198 it imposes large computational demands per iteration, making it computationally ineffective. Therefore, it is important to
199 select appropriate algorithms that effectively find the important regions and converge fast.

200 **VEGAS Formulation for PSHA** VEGAS is a non-parametric AIS algorithm that iteratively identifies the optimal proposal
201 density, q^* (Lepage, 1978, 2021). The algorithm has been developed and widely used in computational physics (Kersevan
202 and Richter-Was, 2013; Alwall et al., 2014), and is currently applied to chemistry, astrophysics, finance, and medical statistics
203 (Campolieti and Makarov, 2007; Garberoglio and Harvey, 2011; Ray et al., 2011; Sanders, 2014).

204 The VEGAS algorithm is conceptually straightforward and is recognized for its rapid convergence, especially when the
205 random variables involved are independent (Lepage, 1978, 2021). In PSHA, the variable \mathcal{E} is always considered an indepen-
206 dent variable (Eq. (3)). Additionally, the variables M and R are also treated as independent under point source assumption
207 (Eq. (4)). Notice, however, that when the finite-fault rupture model, in which the rupture dimension changes with magni-
208 tude, is adopted, M and R can be correlated, and the distribution of R is conditional on the magnitude M . Also, note that the
209 estimate remains unbiased in this case, because its calculation is still within the IS framework.

210 In three-dimensional integration, which is the case of PSHA, VEGAS employs N^3 cuboids that are independently parti-
 211 tioned. The probability assigned to each cuboid and the total number of partitioned cuboids (N^3) is preserved across the
 212 iteration steps. However, the IS sampling density changes because the algorithm updates the cuboid's size depending on
 213 its contribution to the integration. If a cuboid's contribution is low, its size grows in the next step, lowering its probability
 214 density. Conversely, when a cuboid's contribution is high, it shrinks in the following step, elevating the probability density.
 215 Ideally, when every cuboid's contribution to the integration becomes identical, we find the proposed optimal density, q^* , and
 216 the algorithm is terminated. The framework to find optimal IS densities for PSHA using the VEGAS algorithm is illustrated
 217 with a simple point seismic source example in the following paragraphs.

218 First, we adopt an independently distributed joint probability function as IS density, significantly improving the efficiency
 219 of the iterative updating process in multi-dimensional PSHA integration. Thus, $q_{M,R,\mathcal{E}}(m, r, \varepsilon) = q_M(m)q_R(r)q_{\mathcal{E}}(\varepsilon)$. Then, the
 220 integration ranges for each random variable— M , R , and \mathcal{E} —are divided into N grids with the same volume. This division is
 221 designed to generate cuboids of constant probability:

$$\begin{aligned}
 M &: m_{i-1} \leq M < m_i \quad (i = 0, 1, 2, \dots, N), \Delta m_i = m_i - m_{i-1} \\
 R &: r_{j-1} \leq R < r_j \quad (j = 0, 1, 2, \dots, N), \Delta r_j = r_j - r_{j-1} \\
 \mathcal{E} &: \varepsilon_{k-1} \leq \mathcal{E} < \varepsilon_k \quad (k = 0, 1, 2, \dots, N), \Delta \varepsilon_k = \varepsilon_k - \varepsilon_{k-1}
 \end{aligned} \tag{20}$$

222 The number of grids, N , is chosen to be 50 as suggested by [Lepage \(1978\)](#) and because we consider this value makes grids
 223 sufficiently fine to capture the actual distribution of the optimal density q^* . The probability for each cuboid from Eq. (20) is
 224 set to $1/N^3$ to ensure that the initial IS density function is uniformly distributed across the entire domain. As a result, the
 225 initial probability density $q^{(0)}$ for a specific cuboid is

$$\begin{aligned}
 q^{(0)}(m_{i-1} \leq M < m_i, r_{j-1} \leq R < r_j, \varepsilon_{k-1} \leq \mathcal{E} < \varepsilon_k) &= q_M^{(0)}(M) q_R^{(0)}(R) q_{\mathcal{E}}^{(0)}(\mathcal{E}) \\
 &= \frac{1}{N\Delta m_i} \frac{1}{N\Delta r_j} \frac{1}{N\Delta \varepsilon_k}
 \end{aligned}$$

226 Figure 1 illustrates an example of initial partitioning when $N=10$. The gray dots in the figure show an example of MC samples
 227 from this initial q^0 . Note that for easier visualization and understanding, Figure 1 illustrates an example in m - ε , the two-
 228 dimensional space, and not the three-dimensional m , r , and ε space.

229 We update q using MC samples (M, R, \mathcal{E}) to ultimately make it converge to q^* (Eq. (19)). As mentioned earlier, the size
 230 of each cuboid is subject to change while the probability of each cuboid remains constant, i.e., probability density changes.
 231 This adjustment is facilitated through a "subdivision-and-restoration" process ([Lepage, 1978](#)). In this process, each VEGAS
 232 grid is first subdivided into sub-grids given the total number of sub-grids (N_{subgrid}). Since N_{subgrid} should be sufficiently larger
 233 than N to iterate the IS density effectively, we chose N_{subgrid} to be 10,000, which is 200 times greater than N ($=50$). At each

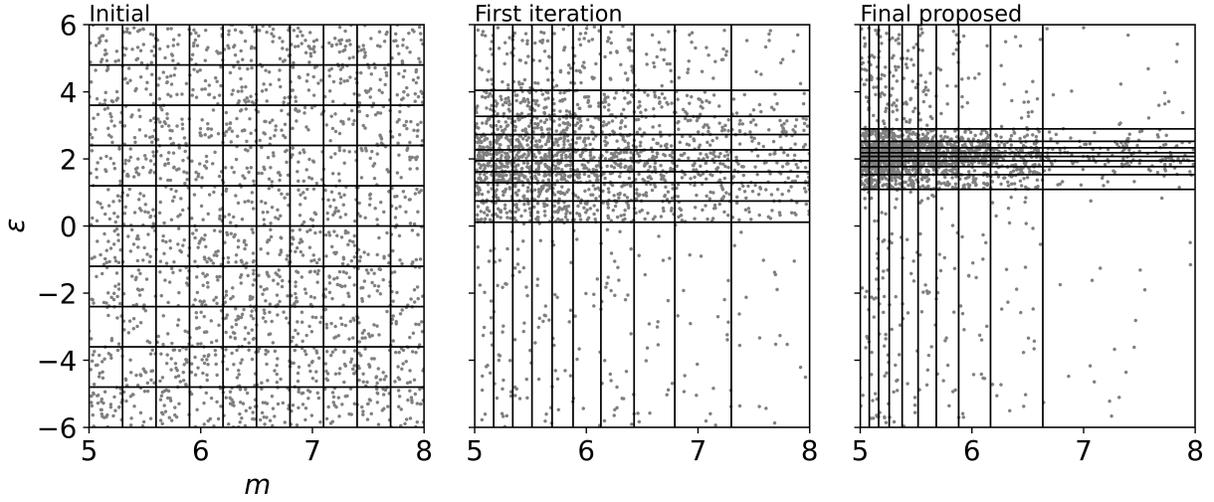


Figure 1. VEGAS iterations of IS density in our AIS PSHA framework. Example for PSHA at PGA = 0.5 g when the site is located 10 km away from a point seismic source ($m_{\min} = 5.0$, $m_{\max} = 8.0$, b -value = 1.0, $\nu = 1.0/\text{yr}$, Ground Motion Model (GMM) = [Sadigh et al. \(1997\)](#)). The total number of MC samples is 2,000, and the number of grids per axis is 10, constituting a total of 100 rectangles (cuboids in actual three-dimensional PSHA integration). The vertical and horizontal black solid lines are boundaries of the m and ε rectangles, and gray dots are the MC samples. Starting from the initial same-size m and ε rectangles, their sizes are adjusted depending on the contribution of each rectangle to the hazard. The final proposed structure of the rectangles gives a highly concentrated probability density at $m \sim 5$ and $\varepsilon \sim 2$. Note that the hazard estimates using initialized density, density after the first iteration, and the final proposed density are 0.0417, 0.0413, and 0.0378, while the true solution is 0.0385.

234 VEGAS grid, the number of sub-grids is proportional to each grid's contribution to the overall integration, d_i :

$$d_i = \frac{\bar{H}_i \Delta x_i}{\sum_i \bar{H}_i \Delta x_i} \quad (21)$$

235 ,where \bar{H}_i in m axis is ([Lepage, 1978](#)):

$$\bar{H}_{i,m} = \sqrt{\sum_{m_{i-1} < M < m_i} \frac{H^2(M, R, \varepsilon; a)}{q_R(R)q_\varepsilon(\varepsilon)}} \quad (22)$$

236 , where $H(M, R, \varepsilon; a)$ is the integrand of PSHA, $I(X > a | M, R, \varepsilon) f_{M,R,\varepsilon}(M, R, \varepsilon)$. Intuitively, \bar{H}_i can be considered as the
 237 marginalization of the overall contribution of the magnitude dimension within the i th grid. Note that the summation of
 238 H^2 s in Eq. (22) is done over all the samples which fall within m_{i-1} and m_i and the division by $q_R q_\varepsilon$ can be considered an
 239 adjustment for unevenly distributed R and ε samples to purely capture the magnitude contribution.

240 Similarly, \bar{H}_i in r and ε axes can be calculated as

$$\bar{H}_{i,r} = \sqrt{\sum_{r_{i-1} < R < r_i} \frac{H^2(M, R, \varepsilon; a)}{q_M(M)q_\varepsilon(\varepsilon)}}$$

241

$$\bar{H}_{i,\varepsilon} = \sqrt{\sum_{\varepsilon_{i-1} < \varepsilon < \varepsilon_i} \frac{H^2(M, R, \varepsilon; a)}{q_R(R)q_M(M)}}$$

242 After the sub-division of each VEGAS grid with the numbers proportional to d_i , the total number of grids is restored to the
 243 original number, N , by merging N_{subgrid}/N consecutive subgrids. The restored grid is the updated density, $q^{(1)}$. The iterative
 244 process described above is repeated until no further changes occur in d_i (Eq. (17)).

245 In the early iterations, researchers have noticed numerical instabilities due to the relatively poor information on the inte-
 246 grand (Lepage, 1978, 2021). However, researchers have also found effective ways to mitigate it through smoothing (Eq. (23))
 247 and damping (Eq. (24)). The smoothed and damped contributions at i th VEGAS grid, d_s and d_d , are calculated as:

$$d_s := \frac{1}{\sum d_i} \times \begin{cases} (7d_i + d_{i+1})/8 & \text{if } i = 0 \\ (d_{i-1} + 6d_i + d_{i+1})/8 & \text{if } i = 1, 2, \dots, N - 1 \\ (d_{i-1} + 7d_i)/8 & \text{if } i = N \end{cases} \quad (23)$$

248

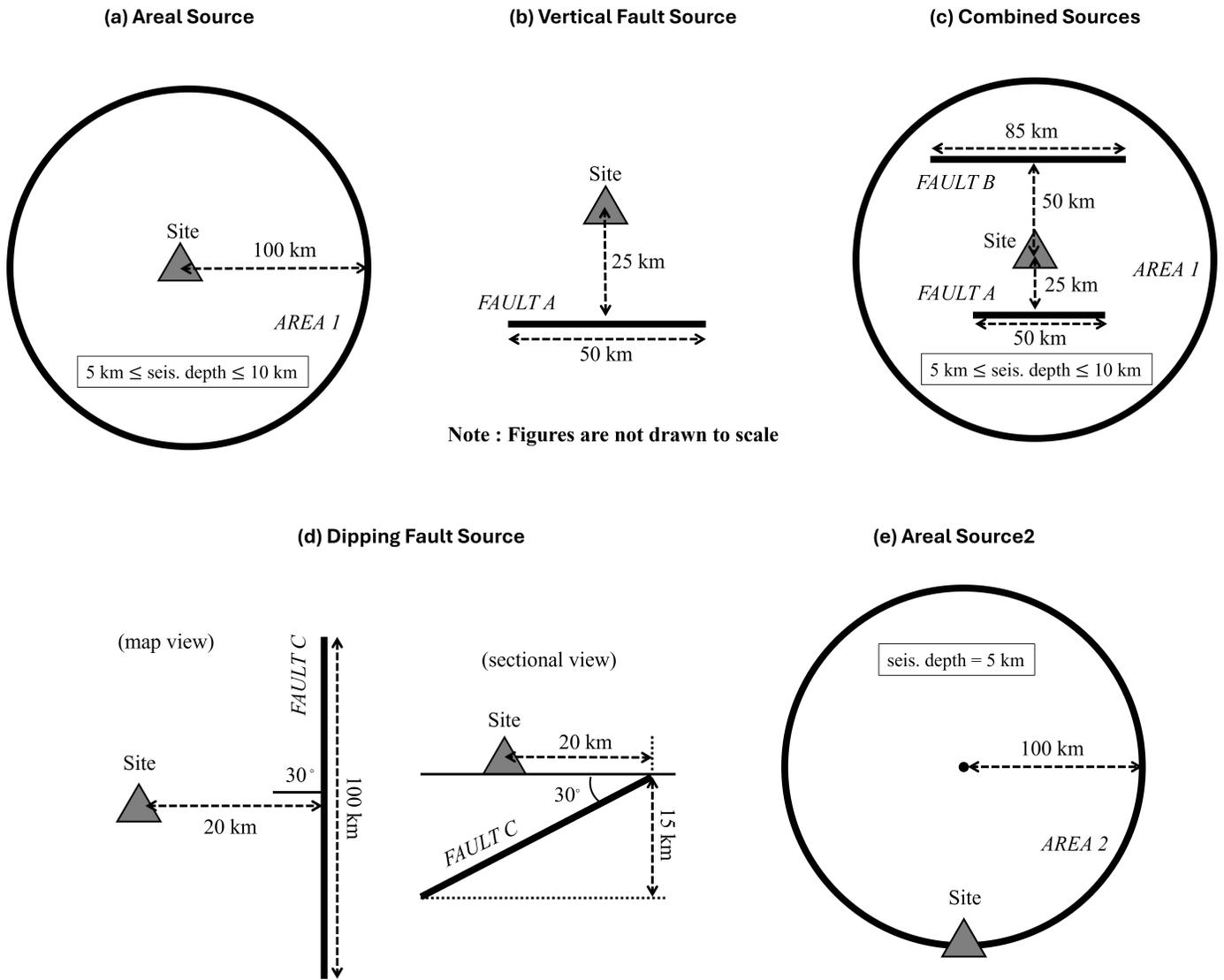
$$d_d := \left(\frac{1 - d_s}{\ln(1/d_s)} \right)^\alpha \quad (24)$$

249 , where α is learning rate. We used α as 1.0 as suggested by Lepage (2021). Through numerical experimenting, we observed
 250 this value allowed most of the PSHA integration to converge within three iterations without causing any numerical
 251 instabilities.

252 This VEGAS iteration process is illustrated in Figure 1. The sizes of the grids are shrunk at $5 < m < 6$ and $\varepsilon \sim 2$, indicating
 253 that the high contributions to the hazard on this range (d_i) in contrast to other less important regions, e.g., at $m > 6.5$ and
 254 $\varepsilon < 1$. Note that the probability of each grid is preserved so that the number of MC samples inside each grid is almost the
 255 same regardless of the grid size. The final proposed density exhibits features that are expected based on intuition. First, the
 256 low contribution of ε less than 0 is understandable. This is because the simulated ground motion intensity, using Sadigh et al.
 257 (1997), never exceeds the target ground motion of 0.5 g at any magnitude, even at $m = 8.0$, when ε is less than 0.1. In addition,
 258 the strong contribution of $m \sim 5$ and $\varepsilon \sim 2.2$ to the integration makes sense because the simulated ground motion intensity
 259 exceeds the target ground motion of 0.5 g when m and ε reach 5 and 2.2, respectively. The decreasing trend of contribution
 260 beyond $m \sim 5$ and $\varepsilon \sim 2.2$ can be interpreted as the exponential decay in the probability of m and ε . This result for a simple
 261 PSHA example implies that the proposed algorithm can be utilized for more complex PSHA integration.

262 NUMERICAL EXAMPLES

263 We tested our proposed AIS PSHA using five numerical examples. For examples 1 to 3, we utilized benchmark problem sets
 264 1.11 and 2.1 from the PEER PSHA code verification project (Hale et al., 2018) to validate the method against the verified



Note : Figures are not drawn to scale

Figure 2. Seismic source geometry for the numerical examples to test our AIS PSHA framework.

265 benchmark PSHA curves (Figure 2 (a)-(c)). The examples encompass various seismic source types such as area, fault, and
 266 combined sources. In Example 4, dipping fault is introduced with widely used model parameters common in PSHA prac-
 267 tice (Figure 2 (d)). For these Examples, the computational efficiency of AIS PSHA is compared with pre-existing numerical
 268 solutions such as Riemann Sum, conventional MC, and IS with uniform sampling distribution (Jayaram and Baker, 2010).
 269 For the accurate Riemann Sum, we followed the procedure of Hale et al. (2018). Finally, in Example 5, we introduced an
 270 advanced version of AIS, termed “smart” AIS, and tested it against a “smart” Riemann Sum variant that uses a strategic grid
 271 spacing to solve problem set 1.10 of Hale et al. (2018) (Figure 2 (e)).

272 For all the examples, the probability of exceedance at eighteen ground motions (PGA), 0.001, 0.01, 0.05, 0.1, 0.15, 0.2, 0.3,
 273 0.35, 0.4, 0.45, 0.5, 0.55, 0.6, 0.7, 0.8, 0.9, and 1.0 g, are estimated. Each example’s model parameters are detailed in Table S1,
 274 and their benchmark hazard curves are shown in Figure 3. The analyses were conducted in Python 3.11 on an Intel Core
 275 i7-13700 2100 MHz processor with 64GB RAM.

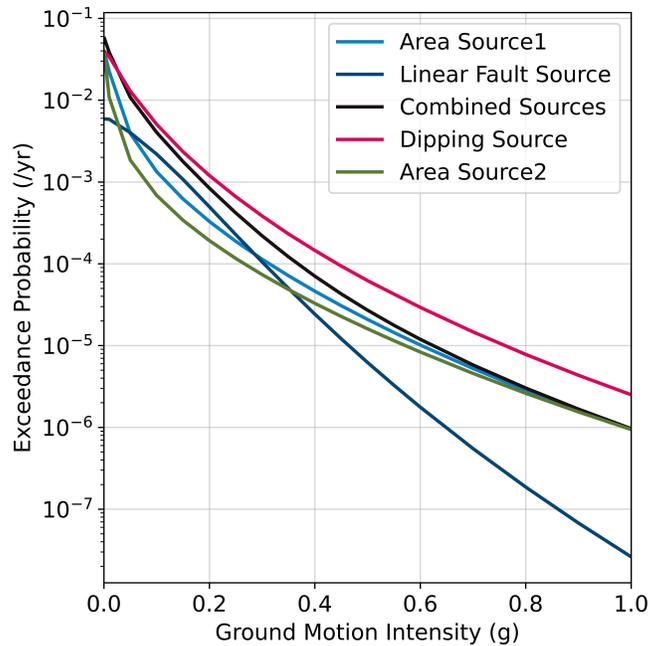


Figure 3. Benchmark PSHA curves for area source 1, linear fault source, combined sources, dipping source, and areal source 2. The benchmark curves are obtained from either Hale et al. (2018) (for areal source 1, combined sources, and areal source 2) or calculation by the authors (linear fault source and dipping source).

Example 1: Areal source

We consider the circular area source with the site at the circle's center (Figure 2). We solved the PSHA using Riemann Sum, conventional MC, uniform IS, and AIS, and show their performance (Figure 4; see Figure S2 for the results on the other ground motion levels). Our numerical experiments show that computational times and COV ("accuracy") are linearly correlated in logarithmic scale, in agreement with the theory (Eq. (11)).

While the computation time and COV vary depending on the target ground motion of interest, our results show our AIS PSHA generally outperforms the other numerical techniques. At a low target ground motion of 0.05g, the computation time to achieve a 1 % COV, i.e., 95 % of the hazard estimates being within 2% error range, is 0.04 seconds for AIS, while it takes 0.20 and 0.39 seconds for conventional MC and IS estimates, respectively, indicating AIS is 4.4 and 8.7 times faster. The computational efficiency becomes extreme at higher ground motions. At 0.3 g, which corresponds to the exceedance probability of $\sim 10^{-4}$ /yr, the computation time to achieve 1 % COV is estimated to be 0.06 seconds for AIS, while it takes 7.14 and 1.81 seconds for conventional MC and IS estimates. In this case, AIS is 111 and 28 times faster than conventional MC and IS, respectively. At the extreme ground motion, 1.0 g, the computation time to achieve 1 % COV is estimated to be 0.09 seconds for AIS, while it takes 633 and 6.53 seconds for conventional MC and IS estimates. In this case, AIS is 7,128 and 73 times faster than conventional MC and IS, respectively. Also, for the 1 % COV case, AIS is $\sim 3.7 \times 10^4$ faster than Riemann Sum. We also note that AIS outperforms IS in all ground motion ranges by a factor of 12 to 73 to achieve a 1 % COV. However, this performance is not always guaranteed because AIS takes t times more computational time than IS with the same N due to the t iterations

293 to find the optimal IS density. Thus, our findings imply that AIS PSHA with the VEGAS algorithm can find (close to) optimal
294 IS density quickly.

295 It is also noteworthy that with the fixed number of samples (i.e., constant computational burden), AIS's estimation accu-
296 racy is quite similar across different target ground motions, while conventional MC's accuracy decreases sharply for higher
297 ground motions (Figure 5). According to Eq. (10), COV of MC hazard estimates for the ground motion corresponding to
298 10^{-6} /year will increase to 10% if it initially was 1 % for the ground motion corresponding to 10^{-4} /year. From our numerical
299 experiments, we found errors grow from 0.54 to 5.60 % with $N_s = 10^7$ MC samples for these ground motions (0.3g and 1.0g).
300 In contrast, for a similar computation time, we found AIS had errors ranging from 0.51 to 0.67 % for these quite different
301 ground motion levels (Figure 5).

302 This finding is key for PSHA as the computational bottleneck is at the highest ground motion intensity. Using conventional
303 MC PSHA, the hazard analyst has no choice but to largely increase the number of samples to estimate hazard accurately
304 at high ground motions even though such a large number would not be necessary for low ground motions. This makes
305 the conventional MC highly inefficient, and consequently, the efficiency of the conventional MC for lower ground motions
306 cannot be considered a real advantage for PSHA. AIS PSHA overcomes this problem by adopting different optimal densities
307 at different ground motion intensities, making the computational burden to achieve the similar estimation error almost flat
308 for any ground motion intensity, as shown in Figure S3. We observed that that AIS PSHA achieves 2.5 % and 1.0 % COV with
309 $N_s \sim 10,000$ and 50,000, respectively (Figure S3).

310 Another key advantage of our proposed AIS PSHA is the co-production of disaggregation curves at no extra computational
311 cost. From Eq. (19), optimal IS densities are theoretically equivalent to hazard disaggregation. We showed that the benchmark
312 marginal distributions of hazard disaggregation obtained from Riemann Sum closely match the iterated IS density from AIS
313 PSHA (Figure 6; see Figure S4 for the results on the other ground motion levels). We used the Kolmogorov-Smirnov (K-S)
314 D statistic (Kolmogorov, 1933) to quantify their similarities. K-S D statistic measures the maximum difference between two
315 cumulative distribution functions (CDF):

$$D = \max_x |F(x) - Q(x)|$$

316 , where x is m , r , or ε , $F(x)$ is the marginal CDF of disaggregation result, $Q(x)$ is the CDF of proposed IS density obtained from
317 VEGAS. D is larger for more dissimilar CDFs, with a maximum possible value of one. Conversely, if two CDFs are identical,
318 D is zero. Thus, D close to zero indicates that the two probability distributions are similar. For reference, examples of K-S D
319 statistics are illustrated in Figure S1.

320 We calculated D for m , r , and ε at all the ground motion intensities (Figure S5). We found the maximum D values for
321 m , r , ε were 0.032, 0.113, and 0.092, respectively, and the minimum values were 0.019, 0.026, and 0.017, indicating a strong
322 resemblance between the two distributions.

323 We also estimated the differences between the disaggregation's mean values from the proposed IS density and Riemann
324 Sum (Figure S5). Note that the disaggregation distributions and proposed IS densities vary for different ground motion levels.
325 Thus, their mean values also vary. We found the maximum relative differences were 2.5 %, 22.6 %, and 18.5 % for m , r , ε , and
326 the minimum differences were 1.7 %, 1.0 %, and 4.3 %, respectively. Also, the maximum absolute differences were 0.14, 3.9
327 km, and 0.17 for m , r , ε , and the minimum differences were 0.09, 0.7 km, and 0.0009. The maximum relative difference in
328 distance (r) appears for disaggregation distributions at 0.25 g, where the mean distance obtained from hazard disaggregation
329 is 14.3 km and that from the proposed IS density is 17.5 km. Given that we are typically interested in distance ranges on
330 the order of tens of kilometers (e.g., 0-15 km, 15-25 km, 25-50 km, etc.) rather than a single value (U. S. Nuclear Regulatory
331 Commission, 2007), this difference is not crucial in determining the controlling earthquake for critical infrastructures. In
332 addition, IS densities still show small K-S D statistics and their mode almost matches each other, as shown in Figure 6.

333 **Example 2: Fault source**

334 We consider a 50 km-length vertical fault 25 km away from the site (Figure 2 (b)). We adopted a finite-dimension rupture
335 model, which results in a distance distribution dependent on magnitude. We found this dependency diminishes the perfor-
336 mance of VEGAS AIS because the VEGAS algorithm assumes the independently distributed random variables. For example,
337 for a ground motion intensity of 0.05 g, AIS is slower than conventional MC (Figure S6), while AIS outperformed conven-
338 tional MC at the same ground motion intensity when point source assumption was made, as seen earlier. However, as the
339 ground motion intensity increases, the computational gap between the two methods becomes smaller rapidly and closes at
340 0.2g. For higher ground motions, AIS outperforms conventional MC. At 0.3 g, corresponding to $\sim 10^{-4}$ exceedance prob-
341 ability, for a 1% COV, AIS, IS, and conventional MC take 0.25, 71, and 0.43 seconds, respectively, i.e., AIS MC is 277 and
342 1.7 times faster. At more extreme ground motion (1.0 g), AIS, IS, and conventional MC take 0.25, 96, and 4,294 seconds,
343 respectively, i.e., AIS is 390 and 17,448 times faster (Figure 4; see Figure S6 for the results on the other ground motion levels).
344 We observed that AIS PSHA estimates approximate the true hazard curve at 2.5 % and 1.0 % COV when N_s is $\sim 50,000$ and
345 300,000, respectively (Figure S7).

346 We also compared hazard disaggregation and the iterated IS density (Figure 6; see Figure S8 for the results on the other
347 ground motion levels) and showed they closely match each other even though the distance distribution depends on the
348 magnitude in this case. It is also noteworthy that the iterated IS density can even reproduce complex densities with disconti-
349 nuousities like the large jump within magnitude distribution (for $m = 6.25$). The K-S D statistic and mean difference of the two
350 distributions are also presented in Figure S9. The maximum values of D in m , r , and ε are 0.30, 0.68, and 0.13, respectively,
351 and the minimum values are 0.02, 0.31, and 0.04. We found that the largest discrepancies occur in the magnitude distribu-
352 tion, but errors can be considered negligible as the mean magnitude difference is within 6 % error. We also note considerable
353 discrepancies in the distance distribution shape (see K-S D statistics of Figure S9) as curves with concentrated probabilities

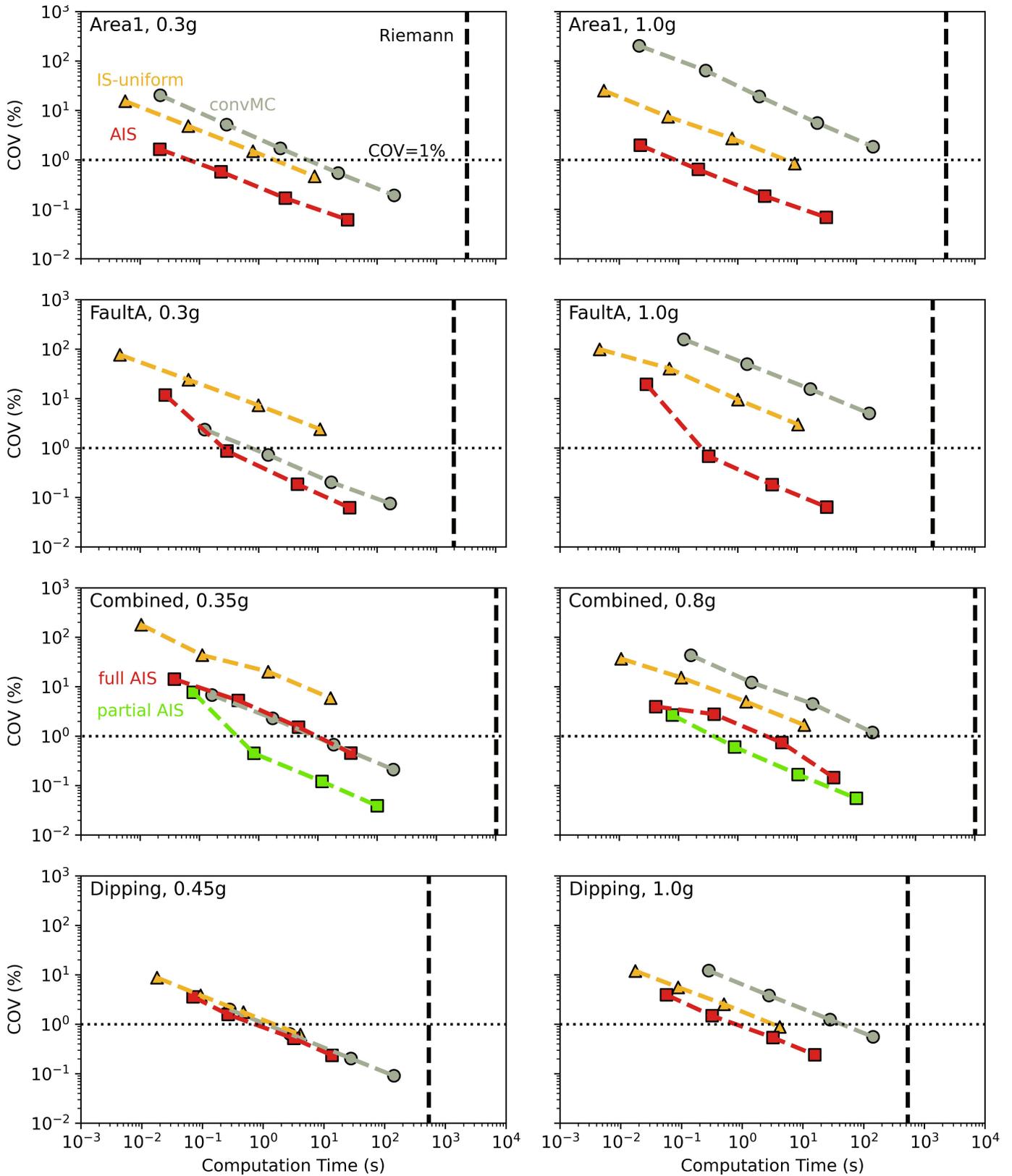


Figure 4. Example 1-4 (Area (Area1), Vertical fault (FaultA), Combined, and Dipping fault (FaultC) sources example): COV as a function of computation time using various PSHA numerical solutions at ground motion intensities that correspond to $10^{-4}/\text{yr}$: 0.3g, 0.3g, 0.35g, and 0.45g for the area, fault, combined, dipping fault sources, respectively, and extreme ground motion intensities: 1.0g ($\sim 10^{-6}/\text{yr}$), 1.0g ($2.62 \times 10^{-8}/\text{yr}$), 0.8g ($3.11 \times 10^{-6}/\text{yr}$), and 1.0g ($2.51 \times 10^{-6}/\text{yr}$) for area, vertical fault, combined, and dipping fault sources, respectively. COV of 1% is denoted as dotted horizontal lines. As the target ground motion increases, the performance of AIS makes a dramatic improvement in terms of both accuracy and computational cost.

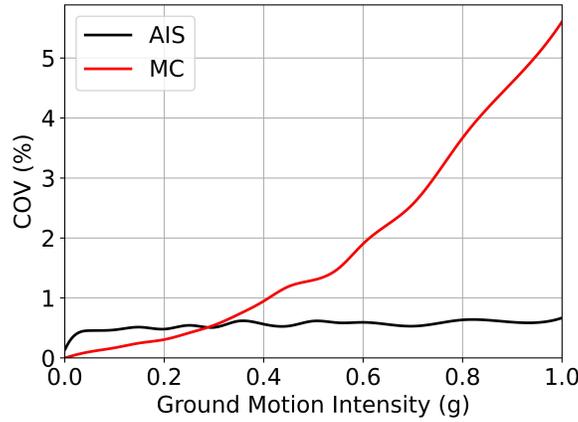


Figure 5. COV of MC estimates with the ground motion for conventional (red) MC ($N_{s, \text{convMC}} = 10^7$) and AIS (black) with the similar computation time. The COV of error exponentially increases with the ground motion in conventional MC, while that of AIS remains constant.

354 in narrow ranges are harder to estimate for AIS (Figure 6). However, the mean distance is still within a 5 % error range. The
 355 ε is generally in good agreement across all ground motion intensities. We observe higher differences at lower ground motion
 356 because the mean ε is close to zero. For example, the calculated mean ε at ground motion intensity of 0.01 g is 0.022 and
 357 0.007, which is not a large difference in practice.

358 The maximum relative differences in mean values were found to be 6.2 %, 5.3 %, and 129 % for m , r , ε , and the minimum
 359 differences were 0.006 %, 3.9 %, and 2.2 %, respectively. Note that the 129 % of ε case corresponds to the case where the mean
 360 ε is close to zero. The relative difference appears to be slightly higher than the previous areal source example, however, the
 361 absolute difference still remains significantly small; the maximum absolute differences were 0.36, 1.3 km, and 0.20 for m , r ,
 362 ε , and the minimum differences were 0.0004, 1.0 km, and 0.07.

363 **Example 3: Combined sources**

364 In PSHA, we often have multiple seismic sources. We consider one area and two fault sources around the site to represent this
 365 case. This application posits a different mathematical problem than the previous two examples because we must introduce
 366 an additional variable to formulate AIS.

367 First, the probability of earthquake occurrence at i th seismic source can be defined as:

$$P(S = i) = \frac{\nu_i}{\sum_{j=1}^{n_s} \nu_j}$$

368 , where n_s is the number of seismic sources ($n_s = 3$ in this example), and ν_i is the annual earthquake occurrence rate of i th
 369 seismic source. Because the discrete random variables cannot be used in AIS, we define a continuous random variable and
 370 its corresponding probability density function as

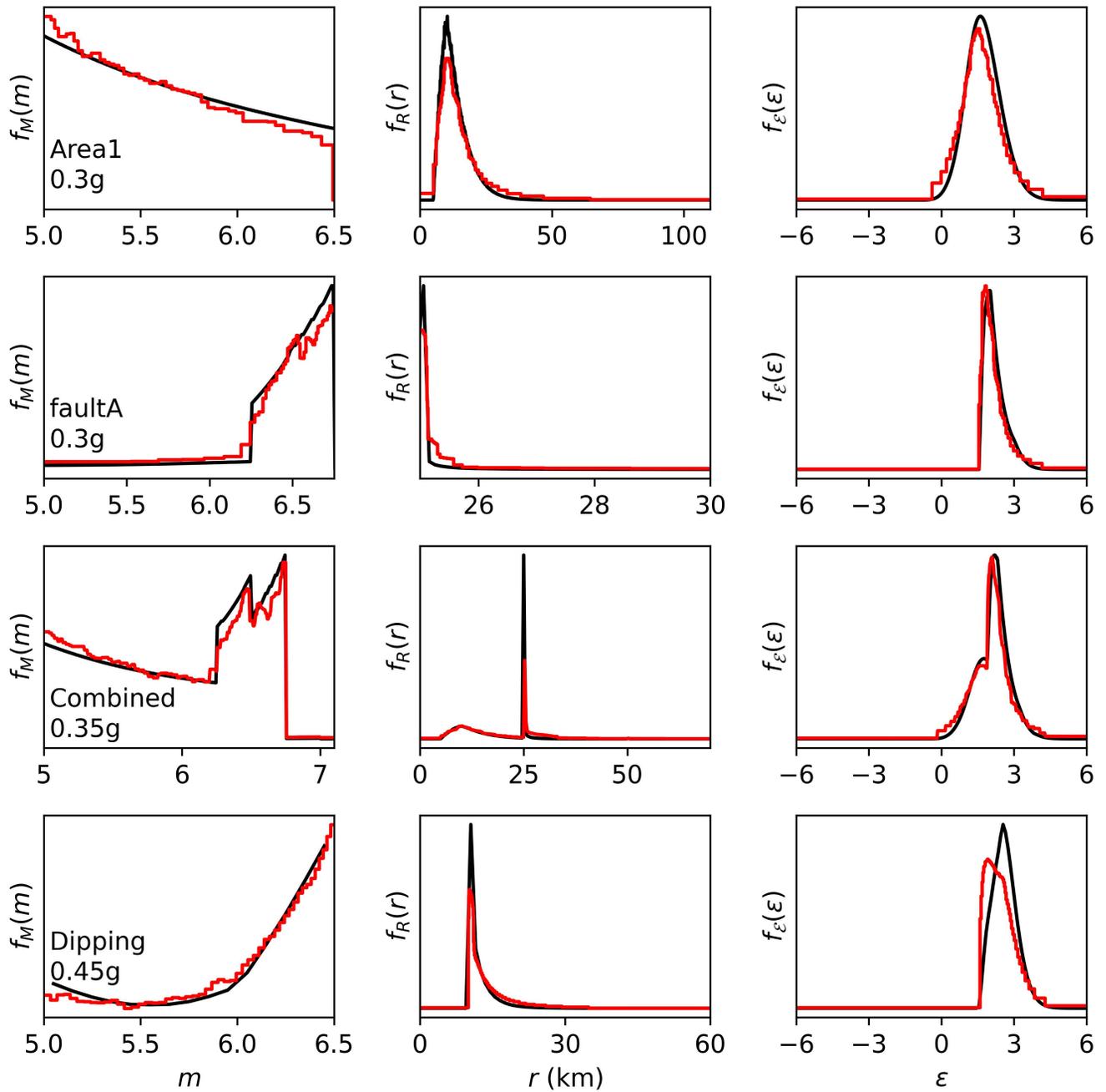


Figure 6. Example 1-4 (Area (Area1), Vertical fault (FaultA), Combined, and Dipping fault (FaultC) sources example): The convergence of m , r , and ε iterated IS densities (red) derived in AIS algorithm to marginal distributions of hazard disaggregation (black).

$$f_s(s) = \begin{cases} 0 & s < 0 \\ P(S=i) & \text{if } i-1 \leq s < i \text{ (} i=1, 2, \dots, n_s \text{)}, \\ 0 & s \geq n_s \end{cases}$$

371 $f_S(s)$ is a piece-wise constant function where the heights are proportional to the corresponding sources' earthquake
 372 occurrence rates. We introduce $f_S(s)$ into the PSHA integration and obtain

$$\Lambda(X > a) = \left(\sum_i^{n_s} \nu_i \right) \int_s \int_\varepsilon \int_r \int_m I(x > a|s, m, r, \varepsilon) f_{M,R,\varepsilon}(m, r, \varepsilon|s) f_S(s) dm dr d\varepsilon ds. \quad (25)$$

373 Solving Eq. (25) posits computational challenges than the single source problems because $f_S(s)$ have the large jumps at $i =$
 374 $1, 2, \dots, n_s - 1$, and additional dependencies of m, r , and s are introduced. This could further diminish the VEGAS algorithm's
 375 effectiveness. Thus, we test two AIS PSHA approaches for this case: 1) full AIS approach utilizing the Eq. (25) and 2) partial
 376 AIS approach, which is the simple summation of single-source AIS PSHA curves.

377 We compared the computational performance of MC, IS, and full and partial AIS (Figure 4; see Figure S10 for the results on
 378 the other ground motion levels). Like the previous examples, conventional MC is faster for low ground motions, but not high
 379 ones. To achieve COV of 1% at a ground motion of 0.35 g, corresponding to $\sim 1.2 \times 10^{-4}$ annual probability, conventional MC,
 380 IS, and full and partial AIS take 8.33, 891, 10.4, and 0.43 seconds, respectively, i.e., partial AIS is the most efficient algorithm,
 381 $\sim 19, 2,087$, and 24 times faster than conventional MC, uniform IS, and full AIS, respectively. At more extreme 0.8 g case,
 382 conventional MC, IS, and full and partial AIS take 295, 49, 3.04, and 0.39 seconds, respectively; partial AIS is $\sim 756, 127$, and
 383 8 times faster than conventional MC, uniform IS, and full AIS, respectively.

384 Figures S11 and S12 present the accuracy of both AIS approaches with different sample sizes. The contrast between Figures
 385 S11 (c) S12 (c) shows that partial AIS has smaller error than full AIS even with fewer samples ($N_s, 150,000 < 500,000$). For
 386 partial AIS, it estimates the true hazard curve at 2.5% and 1.0% COV when N_s is $\sim 150,000$ and 300,000, respectively (Figure
 387 S12).

388 The hazard disaggregation results, compared with the proposed IS density obtained from partial AIS PSHA, are also shown
 389 in Figure 6 (see Figure S13 for the results on the other ground motion levels). Similar to the previous examples, the proposed
 390 IS density effectively captures the complex features of the disaggregation.

391 **Example 4: Dipping fault source**

392 We also tested our approach on a dipping fault passing beneath the site and GMM of Abrahamson et al. (2014). AIS also
 393 performs efficiently for this example (Figure 4; see Figure S14 for the results on the other ground motion levels). To achieve
 394 a 1% COV at 0.45 g-corresponding to an annual probability of 10^{-4} -AIS is 1.9, 2.4, and 683 times faster than uniform-IS, MC,
 395 and the Riemann Sum, respectively. At 1.0 g, AIS reaches higher efficiency and runs 4, 50, and 633 times faster than uniform-
 396 IS, MC, and the Riemann Sum, respectively. A comparison between the proposed IS density and the hazard disaggregation
 397 is also illustrated in Figure 6. Results for the other ground motion levels are available in Figure S15. Both generally exhibit
 398 similar trends, however, some differences observed in this example suggest that obtaining the proposed IS density exactly the
 399 same as the optimal density is a challenging task when dealing with highly correlated (m, r) distribution. Rather, we observe

400 a trade-off between the accuracy of m and ε densities. As the magnitude (m) IS density gives a higher contribution at large
 401 magnitudes than the actual disaggregation, the ε IS density slightly shifted to lower values than the actual disaggregation.
 402 Although differences are apparent in this example, it is important to note that we employed an extremely fine discretization
 403 for the comparison. This discrepancy may not be significant when using the coarser grids typically adopted for engineering
 404 applications (U. S. Nuclear Regulatory Commission, 2007; Petersen et al., 2024).

405 **Example 5: Smart Riemann Sum versus Smart AIS PSHA**

406 In the previous sections, we compared the efficiency and accuracy of various numerical solutions with AIS PSHA. Here, we
 407 extend this comparison by examining the performance of AIS PSHA against an advanced variant of Reimann sum, sometimes
 408 adopted by PSHA software packages to tackle computational burden (Ordaz et al., 2013). For this comparison, we use an
 409 enhanced version of AIS PSHA that leverages the information from similar ground motions to reduce computational times
 410 that otherwise would linearly increase with the number of ground motions of interest (N_a).

411 **Smart Riemann Sum** The smart Riemann Sum essentially follows the same approach as Eq. (8) but incorporates three
 412 primary modifications.

413 First, the summation over the distance, \sum_r , is replaced by a summation over rupture locations (latitude, longitude, and
 414 depth), $\sum_\phi \sum_\psi \sum_z$, accounting for the variety of distance metrics used in modern GMMs, such as R_{rup} , R_{JB} , R_x , etc. Second,
 415 $I(\cdot)$ is replaced by $P(X > a | m, \phi, \psi, z)$, which represents the probability that simulated ground motion exceeds the pre-defined
 416 threshold, a . $P(\cdot)$ can be computed using a pre-computed table of normal distribution CDF or rational Chebyshev approxi-
 417 mation algorithm (Cody, 1969), significantly reducing the computational burden compared to brute-force integration over ε .
 418 The third modification involves the use of varying bin sizes in the spatial domain, or a smart-grid. Finer bins are used closer
 419 to the site, while coarser bins are used at greater distances owing to the reduction in hazard sensitivity to changes in rupture
 420 location at farther distances. A common strategy for this is logarithmic uniform spacing (Appendix C). Incorporating these
 421 three modifications results in a revised version of the Riemann Sum PSHA, as shown below:

$$\lambda(X > a) = \nu \sum_{l=1}^{N_z} \sum_{k=1}^{N_\psi} \sum_{j=1}^{N_\phi} \sum_{i=1}^{N_m} P(X > a | m_i, \phi_j, \psi_k, z_l) f_{M,\Phi,\Psi,Z}(m_i, \phi_j, \psi_k, z_l) \Delta m \Delta \phi_j(r) \Delta \psi_k(r) \Delta z \quad (26)$$

422 **Smart AIS PSHA** The AIS PSHA approach presented earlier uses a uniform distribution as the initial IS density. With the
 423 smart AIS PSHA, however, we compute the hazard sequentially from the lowest to the highest ground motion levels. Thus,
 424 we can utilize the iterated IS density obtained at a previous ground motion intensity as the initial IS density for the next
 425 (slightly higher) ground motion, e.g., approximated q^* for the hazard rate at 0.2g as initial IS density for 0.3g. While this
 426 approach cannot be applied at the lowest ground motion intensity, it is worth noting that determining the optimal density
 427 at extremely low ground motion levels is a trivial problem as $I(\cdot) \equiv 1$ (Eq. (19)). This smart approach reduces the number

428 of iterations required to approximate the optimal density, further enhancing computational efficiency while maintaining its
429 robustness.

430 **Numerical Example** In this section, we compare the efficiency and accuracy of the smart Riemann Sum, (naive) AIS, and
431 the smart AIS using PSHA test problem 1.10 from Hale et al. (2018) (Fig. 2 (e)). For this example, the low probability hazard
432 is reported to be highly sensitive to the starting point of the Riemann-sum grids since the site only has the sources on one
433 side.

434 Like naive Riemann Sum, the efficiency of the smart Riemann Sum is dependent on the spatial grid spacing, with a trade-
435 off between the computational efficiency and estimation accuracy. To estimate COV of the smart Riemann Sum at a given
436 grid structure, we assess the variability of the hazard estimates with respect to the randomized locations within each grid
437 cell, as the error metric for the smart Riemann Sum at a given grid structure.

438 For spatial grid spacing of smart Riemann Sum, we conducted multiple sensitivity tests and determined that a magnitude
439 bin size of 0.1 is sufficiently small. With this magnitude bin size, we evaluated the hazard computation time and COV for
440 varying spatial log-uniform bin sizes (δ in Appendix C) of 0.01, 0.02, 0.03, 0.04, 0.05, 0.07, 0.1, and 0.2.

441 Our findings indicate that AIS outperforms the smart Riemann Sum, consistent with previous numerical examples, par-
442 ticularly in estimating low-probability hazards (Figure 7; see Figure S16 for the results on the other ground motion levels
443). For example, at 1.0 g ($\sim 10^{-6}$ /year), achieving a COV of 1% requires computation times that are 43 and 130 times faster
444 with naive AIS and smart AIS, respectively, compared to the smart Riemann Sum. At lower ground motions (e.g., 0.001 g),
445 the Riemann Sum proves to be more efficient than AIS (Figure S16). It is important to note that the Riemann Sum uses the
446 same grid spacing for all ground motion intensities to eliminate the need for repeated ground motion simulations. Therefore,
447 to ensure hazard accuracy across all ground motion levels, it is fair to compare the Riemann Sum based on its worst-case
448 performance, i.e., at 1.0 g.

449 Here, we note that identifying the “optimal” grid settings for Riemann Sum necessitates many sensitivity analyses based
450 on grid size, introducing an additional computational burden, and our comparison does not account for the time required for
451 these analyses. On the other hand, AIS allows for error estimation using Eq. (18) to compute the error of AIS hazard without
452 running the code repeatedly, which is also one of the benefits of using AIS.

453 We also observed that the accuracy of the Riemann Sum saturates at different levels for various ground motion intensities
454 (Figure 7). This residual error appears to be from coarse gridding at greater distances inherent in log-uniform spacing. While
455 log-uniform spacing significantly enhances the efficiency of the Riemann Sum, it introduces irreducible errors that cannot
456 be mitigated. In contrast, with AIS and smart AIS, increasing the number of samples, proportional to the computation time,
457 leads to an exponential reduction in error.

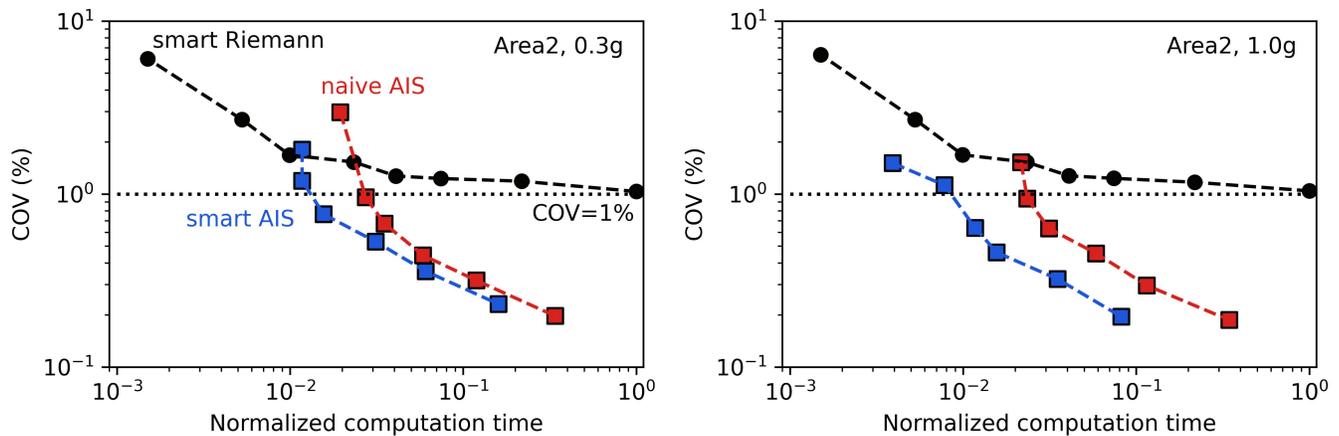


Figure 7. COV of smart Riemann Sum (black), AIS (red), and smart AIS (blue) hazard estimates as functions of computation time at ground motions of 0.3 g and 1.0 g, corresponding to $\sim 10^{-4}/\text{yr}$ and $10^{-6}/\text{yr}$, respectively.

AIS PSHA: POTENTIAL IMPACT ON PRACTICAL APPLICATION

We tested the AIS PSHA using various numerical examples; however, there are many more complex cases encountered in practice. In this section, we discuss how PSHA model complexity might affect the computational efficiency of AIS PSHA from two perspectives: (1) the propagation of epistemic uncertainty and (2) the use of more sophisticated seismic source and ground motion models.

In practice, researchers typically set multiple models and parameters to propagate the epistemic uncertainty of seismic hazards by organizing these into a logic-tree structure. The total computation time required to obtain the mean hazard and its percentiles is the product of the time needed to compute an individual hazard curve and the number of logic-tree end branches. Thus, any reduction in computation time for a single PSHA curve using our proposed approach also proportionally reduces the total computation time for calculating the mean and percentiles. This paper also demonstrated how to incorporate discrete probability distribution into the AIS framework in the combined sources example. This approach can similarly be applied to logic trees to model groups of branches simultaneously. However, readers should note that the full AIS strategy, which incorporates all elements into the VEGAS AIS framework, may sometimes be less efficient than calculating individual hazard curves separately and summing them, the partial AIS strategy, due to the increased dependencies among variables. While we could modify the algorithm to account for variable dependencies at an additional computational cost, we consider this an exciting area for future research beyond the scope of this study.

Furthermore, modern seismic source models are becoming more complex as knowledge of seismogenic faults accumulates (e.g., UCERF3; Field et al. (2014)). These complex models may introduce additional dependencies among variables, which could reduce the efficiency of VEGAS AIS, as demonstrated in the floating rupture fault source examples (numerical examples 2 and 4). However, the dependency of the distance distribution given magnitude decreases rapidly as the distance to the

478 source grows, implying that m and r can be treated as independent. In these cases, VEGAS AIS remains highly efficient. In
479 fact, our examples represent worst-case scenarios in terms of AIS efficiency since the site is located close to the fault.

480 Similarly, modern ground motion models (GMMs) can increase the computational burden of PSHA. Recent empirical
481 GMMs incorporate numerous additional terms to capture physical mechanisms not considered before (e.g., directivity effects,
482 hanging-wall effects, depth to bedrock), increasing the number of coefficients by a factor of ~ 10 (Bommer et al., 2010). In
483 some recent PSHA projects, researchers have adopted physics-based ground motion simulations (Milner et al., 2021), fur-
484 ther raising computational demands. AIS can mitigate the increased computational burden associated with complex ground
485 motion models, as AIS requires fewer ground motion samples than traditional numerical methods. Moreover, the compu-
486 tation time for updating the IS density—a unique feature of AIS—remains constant regardless of model complexity (Figure
487 S17).

488 CONCLUSION

489 We proposed a novel computational framework for PSHA based on an implementation of the VEGAS algorithm. Through
490 comprehensive testbeds, we investigated the computational performance of this new approach, covering widely adopted
491 source types in PSHA practice, including area, vertical and dipping fault, and combined sources.

492 We compared the proposed method to three existing computational frameworks: a) Riemann Sum, which has exponen-
493 tially increasing computational costs for finer grid sizes; b) conventional MC, which requires a substantially long catalog
494 with synthetic earthquakes, particularly for large seismic intensities; and c) importance sampling (IS) with simple (uniform)
495 IS distributions.

496 Our findings indicate that AIS PSHA outperforms all other computational frameworks. AIS PSHA can dramatically reduce
497 computational times by factors up to $\sim 3.7 \times 10^4$ compared to traditional Riemann Sum. AIS PSHA was also 7.1×10^3 faster
498 than the conventional MC while maintaining a 1% COV. Additionally, AIS PSHA was up to 70 times faster than IS PSHA,
499 demonstrating that the VEGAS algorithm can approximate optimal IS distributions quickly and well.

500 We showed that AIS PSHA requires a similar computation time for any ground motion, making its application to larger
501 ground motions with low probability substantially more efficient than conventional MC. In PSHA practice, the computa-
502 tional demands are dominated by large ground motions that need substantially more samples than lower ground motions.
503 In contrast, AIS PSHA only requires a similarly low number of samples for all ground motion levels because it finds optimized
504 IS distributions for each.

505 We also showed that AIS PSHA finds approximated disaggregation curves at no extra computational cost based on theo-
506 retical insights showing that optimal IS densities are equivalent to disaggregation distributions. We showed empirically that
507 the hazard disaggregation and iterated IS densities from AIS PSHA are fairly similar by comparing the statistical properties
508 of the two distributions, e.g., K-S D statistics < 0.113 and mean values differences of $< 4.3\%$.

509 Our study indicated that our AIS PSHA implementation works extremely well for point sources, where the magnitude
510 and distance are independent random variables. We also applied the algorithm to problems employing the finite-rupture
511 model, and it still outperformed the pre-existing algorithms up to by factors of >2,000. It appears to be less effective than
512 the point source example because it introduces a dependency between the magnitude and distance. Our implementation of
513 the VEGAS algorithm uses independent random variables for the IS distribution, thus making it less effective for the finite-
514 rupture case. However, future implementations can also consider other versions of AIS algorithms where the variables are
515 correlated; though computational demands for the AIS iteration would increase.

516 In the case of combined seismic sources, we proposed two strategies: 1) incorporating the source random variable into the
517 AIS PSHA framework and 2) simple summation of AIS PSHA curves for individual sources. Both strategies outperformed
518 the traditional methods up to by a factor of ~ 580 . However, the second strategy was more efficient than the first one by a
519 factor of 6. The first strategy added dependencies to the seismic hazard distribution, making the VEGAS less effective.

520 Additionally, we demonstrated that the algorithm can be even more efficient through a “smart” AIS, which uses the ini-
521 tial guess as the proposed IS density from adjacent ground motion intensity. Our findings indicate that this “smart” AIS
522 demonstrates up to a 130-fold improvement in efficiency compared to the Riemann Sum with optimized spatial grid spacing.

523 In sum, AIS can be applied to any PSHA computation, leading us to expect widespread application of the method.
524 Specifically, we consider our proposed AIS PSHA to be significantly beneficial for large-scale projects that involve numerous
525 logic tree branches and have extreme computational demands.

526 DATA AND RESOURCES

527 The source code for computing PSHA using the framework explained in this paper is available at
528 https://github.com/sehoung/ais_psha. Figures were created using Matplotlib (Hunter, 2007) and Microsoft PowerPoint
529 (<http://office.microsoft.com>; last accessed June 2024). Supplemental Material for this article includes one Text, one
530 Algorithm, one Table, and 17 Figures: Text S1 presents an explanation of the VEGAS AIS algorithm, Algorithm S1 shows
531 the pseudocode of VEGAS AIS PSHA, Table S1 shows model parameters of the numerical examples, Figure S1 shows the
532 examples of K-S D statistics, Figures S2, S6, S10, S14, and S16 show COV as a function of computation time of various
533 numerical solutions, Figures S3, S7, S11, and S12 show the box plot of AIS PSHA estimates, Figures S4, S8, S13, S15 show
534 the convergence of iterated IS densities to the hazard disaggregation, Figures S5 and S9 present the K-S D statistic and mean
535 difference between hazard disaggregation and the iterated IS density, and Figure S17 presents the contribution of each part
536 of VEGAS AIS to the total computation time.

537 **DECLARATION OF COMPETING INTERESTS**

538 The authors have filed a patent application related to the method discussed in this paper. This application is relevant to the
539 research presented and could represent a potential financial interest.

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673 **LIST OF FIGURE CAPTIONS**

- 674 • **Figure 1.** VEGAS iterations of IS density in our AIS PSHA framework. Example for PSHA at $PGA = 0.5$ g when the site
675 is located 10 km away from a point seismic source ($m_{\min} = 5.0$, $m_{\max} = 8.0$, b -value = 1.0, $\nu = 1.0/\text{yr}$, Ground Motion
676 Model (GMM) = Sadigh et al. (1997)). The total number of MC samples is 2,000, and the number of grids per axis is 10,
677 constituting a total of 100 rectangles (cuboids in actual three-dimensional PSHA integration). The vertical and horizontal
678 black solid lines are boundaries of the m and ϵ rectangles, and gray dots are the MC samples. Starting from the initial
679 same-size m and ϵ rectangles, their sizes are adjusted depending on the contribution of each rectangle to the hazard. The
680 final proposed structure of the rectangles gives a highly concentrated probability density at $m \sim 5$ and $\epsilon \sim 2$. Note that
681 the hazard estimates using initialized density, density after the first iteration, and the final proposed density are 0.0417,
682 0.0413, and 0.0378, while the true solution is 0.0385.
- 683 • **Figure 2.** Seismic source geometry for the numerical examples to test our AIS PSHA framework.
- 684 • **Figure 3.** Benchmark PSHA curves for area source 1, linear fault source, combined sources, dipping source, and areal
685 source 2. The benchmark curves are obtained from either Hale et al. (2018) (for areal source 1, combined sources, and
686 areal source 2) or calculation by the authors (linear fault source and dipping source).
- 687 • **Figure 4.** Example 1-4 (Area (Area1), Vertical fault (FaultA), Combined, and Dipping fault (FaultC) sources example):
688 COV as a function of computation time using various PSHA numerical solutions at ground motion intensities that cor-
689 respond to $10^{-4}/\text{yr}$: 0.3g, 0.3g, 0.35g, and 0.45g for the area, fault, combined, dipping fault sources, respectively, and
690 extreme ground motion intensities: 1.0g ($\sim 10^{-6}/\text{yr}$), 1.0g ($2.62 \times 10^{-8}/\text{yr}$), 0.8g ($3.11 \times 10^{-6}/\text{yr}$), and 1.0g ($2.51 \times 10^{-6}/\text{yr}$) for
691 area, vertical fault, combined, and dipping fault sources, respectively. COV of 1% is denoted as dotted horizontal lines. As
692 the target ground motion increases, the performance of AIS makes a dramatic improvement in terms of both accuracy
693 and computational cost.
- 694 • **Figure 5.** COV of MC estimates with the ground motion for conventional (red) MC ($N_{s, \text{convMC}} = 10^7$) and AIS (black)
695 with the similar computation time. The COV of error exponentially increases with the ground motion in conventional
696 MC, while that of AIS remains constant.
- 697 • **Figure 6.** Example 1-4 (Area (Area1), Vertical fault (FaultA), Combined, and Dipping fault (FaultC) sources example):
698 The convergence of m , r , and ϵ iterated IS densities (red) derived in AIS algorithm to marginal distributions of hazard
699 disaggregation (black).
- 700 • **Figure 7.** COV of smart Riemann Sum (black), AIS (red), and smart AIS (blue) hazard estimates as functions of
701 computation time at ground motions of 0.3 g and 1.0 g, corresponding to $\sim 10^{-4}/\text{yr}$ and $10^{-6}/\text{yr}$, respectively.

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714 The mean of MC hazard estimates, $\hat{\lambda}$, $E[\hat{\lambda}]$, can be derived as follows:

$$\begin{aligned}
 E[\hat{\lambda}] &= E \left[\frac{\nu}{N_s} \sum_{i=1}^{N_s} I(X_i > a | M_i, R_i, \mathcal{E}_i) \right] \\
 &= \frac{\nu}{N_s} \sum_{i=1}^{N_s} E [I(X_i > a | M_i, R_i, \mathcal{E}_i)] \\
 &= \frac{1}{N_s} \sum_{i=1}^{N_s} \nu \iiint I(X_i > a | m_i, r_i, \varepsilon_i) f_{M,R,\mathcal{E}}(m_i, r_i, \varepsilon_i) dm_i dr_i d\varepsilon_i \\
 &= \frac{1}{N_s} \sum_{i=1}^{N_s} \lambda(X > a) \\
 &= \lambda(X > a)
 \end{aligned}$$

715 The second moment of $\hat{\lambda}$, $E[\hat{\lambda}^2]$, can be expressed as:

$$\begin{aligned}
 E[\hat{\lambda}^2] &= E \left[\left(\frac{\nu}{N_s} \sum_i^{N_s} I(X_i > a) \right)^2 \right] \\
 &= \frac{\nu^2}{N_s^2} E \left[\left(\sum_i^{N_s} I_i \right)^2 \right]
 \end{aligned}$$

716 , where $I(X_i > a)$ is denoted as I_i for simplicity in the following derivation. The above equation expands as follows:

$$\begin{aligned}
 E[\hat{\lambda}^2] &= \frac{\nu^2}{N_s^2} E [(I_1 + I_2 + \dots + I_{N_s})^2] \\
 &= \frac{\nu^2}{N_s^2} E [I_1^2 + I_2^2 + \dots + I_{N_s}^2 + I_1 I_2 + I_1 I_3 + I_1 I_4 + \dots + I_1 I_{N_s} \\
 &\quad + I_2 I_1 + I_2 I_3 + I_2 I_4 + \dots + I_2 I_{N_s} \\
 &\quad + I_3 I_1 + I_3 I_2 + I_3 I_4 + \dots + I_3 I_{N_s} \\
 &\quad \vdots \\
 &\quad + I_{N_s} I_1 + I_{N_s} I_2 + I_{N_s} I_3 + \dots + I_{N_s} I_{N_s-1}] \\
 &= \frac{\nu^2}{N_s^2} E \left[\sum_i^{N_s} I_i^2 + \sum_{i \neq j}^{N_s^2 - N_s} I_i I_j \right]
 \end{aligned}$$

717 Here, I_i^2 is the same as I_i since I_i^2 also takes one if and only if X_i is greater than a . Also, the expectation operator can go inside

718 the sum terms as it is a linear operator:

$$E[\hat{\lambda}^2] = \frac{\nu}{N_s^2} \sum_i^{N_s} \nu E[I_i] + \frac{1}{N_s^2} \sum_{i \neq j}^{N_s^2 - N_s} \nu^2 E[I_i I_j]$$

719 Here, $\nu E[I_i]$ is equal to λ . Also, $I_i I_j$ takes one if and only if both X_i and X_j exceed a . Given that the ground motion samples
 720 (X_i and X_j) are extracted independently, $\nu^2 E[I_i I_j]$ is identical to $\nu E[I_i] \nu E[I_j]$, which is λ^2 :

$$\begin{aligned} E[\hat{\lambda}^2] &= \frac{\nu}{N_s^2} \sum_i^{N_s} \lambda + \frac{1}{N_s^2} \sum_{i \neq j}^{N_s^2 - N_s} \lambda^2 \\ &= \frac{\nu \lambda - \lambda^2}{N_s} + \lambda^2 \end{aligned}$$

721 Hence,

$$\begin{aligned} \text{VAR}[\hat{\lambda}] &= E[\hat{\lambda}^2] - E[\hat{\lambda}]^2 \\ &= E[\hat{\lambda}^2] - \lambda^2 \\ &= \frac{\nu \lambda - \lambda^2}{N_s} \end{aligned}$$

722 APPENDIX B: MEAN AND VARIANCE OF IMPORTANCE SAMPLING PSHA ESTIMATE

723 The mean of IS hazard estimate, $\hat{\lambda}$, $E_q[\hat{\lambda}]$, can be derived as follows:

$$\begin{aligned} E_q[\hat{\lambda}] &= \frac{\nu}{N_s} \sum_{i=1}^{N_s} E_q \left[I(X_i > a | M_i, R_i, \mathcal{E}_i) \frac{f_{M,R,\mathcal{E}}(M_i, R_i, \mathcal{E}_i)}{q_{M,R,\mathcal{E}}(M_i, R_i, \mathcal{E}_i)} \right] \\ &= \frac{1}{N_s} \sum_{i=1}^{N_s} \nu \int I(X > a | m, r, \varepsilon) \frac{f_{M,R,\mathcal{E}}(m, r, \varepsilon)}{q_{M,R,\mathcal{E}}(m, r, \varepsilon)} q_{M,R,\mathcal{E}}(m, r, \varepsilon) dm dr d\varepsilon \\ &= \frac{1}{N_s} \sum_{i=1}^{N_s} \lambda \\ &= \lambda \end{aligned} \tag{27}$$

724 Also, the second moment of $\hat{\lambda}$, $E_q[\hat{\lambda}^2]$, can be expressed as:

$$\begin{aligned} E_q[\hat{\lambda}^2] &= E_q \left[\left(\frac{\nu}{N_s} \sum_i^{N_s} \frac{I(X_i > a) f_{M,R,\mathcal{E}}(M_i, R_i, \mathcal{E}_i)}{q_{M,R,\mathcal{E}}(M_i, R_i, \mathcal{E}_i)} \right)^2 \right] \\ &= \frac{\nu^2}{N_s^2} E_q \left[\left(\sum_i^{N_s} \frac{I_i f_i}{q_i} \right)^2 \right] \end{aligned}$$

725 Note that $I(X_i > a)$, $f_{M,R,\mathcal{E}}(M_i, R_i, \mathcal{E}_i)$, and $q_{M,R,\mathcal{E}}(M_i, R_i, \mathcal{E}_i)$ are denoted as I_i , f_i , and q_i for simplicity in the following
 726 derivation. As derived in Appendix A, the above equation expands as follows:

$$\begin{aligned}
E_q[\hat{\lambda}^2] &= \frac{\nu^2}{N_s^2} E_q \left[\sum_i^{N_s} \left(\frac{I_i f_i}{q_i} \right)^2 + \sum_{i \neq j}^{N_s^2 - N_s} \left(\frac{I_i f_i}{q_i} \right) \left(\frac{I_j f_j}{q_j} \right) \right] \\
&= \frac{\nu^2}{N_s^2} \sum_i^{N_s} E_q \left[\left(\frac{I_i f_i}{q_i} \right)^2 \right] + \frac{\nu^2}{N_s^2} \sum_{i \neq j}^{N_s^2 - N_s} E_q \left[\left(\frac{I_i f_i}{q_i} \right) \left(\frac{I_j f_j}{q_j} \right) \right] \\
&= \frac{\nu^2}{N_s} \frac{1}{N_s} \sum_i^{N_s} E_q \left[\left(\frac{I_i f_i}{q_i} \right)^2 \right] + \frac{1}{N_s^2} \sum_{i \neq j}^{N_s^2 - N_s} \nu E_q \left[\left(\frac{I_i f_i}{q_i} \right) \right] \nu E \left[\left(\frac{I_j f_j}{q_j} \right) \right] \\
&= \frac{\nu^2}{N_s} E_q \left[E_q \left[\left(\frac{I_i f_i}{q_i} \right)^2 \right] \right] + \frac{1}{N_s^2} (N_s^2 - N_s) \lambda^2 \\
&= \frac{\nu^2}{N_s} E_q \left[\left(\frac{I_i f_i}{q_i} \right)^2 \right] - \frac{1}{N_s} \lambda^2 + \lambda^2
\end{aligned}$$

727 Hence,

$$\begin{aligned}
\text{VAR}[\hat{\lambda}] &= E_q[\hat{\lambda}^2] - E_q[\hat{\lambda}]^2 \\
&= \frac{1}{N_s} \left(\nu^2 E_q \left[\left(\frac{I_i f_i}{q_i} \right)^2 \right] - \lambda^2 \right)
\end{aligned}$$

728 APPENDIX C: LOG-UNIFORM SPACING OF SPATIAL GRID

729 By adopting log-uniform spacing in the spatial domain, the i th grid size, Δ_i , can be calculated as:

$$\Delta_i = 10^{\delta \cdot i} (10^\delta - 1)$$

730 , where i is a positive integer increase with the distance from the site, δ is a constant representing the spacing interval in
731 logarithmic scale (e.g., 0.1). Note that simply adopting Δ_i as $10^{\delta \cdot i}$ cannot capture the finer grid spacing at a distance less than
732 one distance unit. The difference between k th and $(k + 1)$ th power of 10 allows us to generate a sufficiently fine grid at those
733 distances. The grid spacing at source location X , r km away from the site location, $\Delta(X; r)$, is as follows:

$$\Delta(X; r) = \Delta_i, \text{ such that } F_{i-1} \leq r < F_i$$

734 , where F_i is the cumulative sum of the grids:

$$F_i = \sum_{k=1}^i \Delta_k, i = 1, 2, 3, 4, \dots$$

735 , where $F_0 = 0$.

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